

Quantisation of constrained systems

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Nonrelativistic harmonic oscillator

Lagrange function of harmonic oscillator:

$$L(x, \dot{x}) = T - V = \frac{m\dot{x}^2}{2} - \frac{1}{2}m\omega^2 x^2$$

Euler–Lagrange equation of motion:

$$\frac{\partial L}{\partial x} - \frac{d}{dt} \frac{\partial L}{\partial \dot{x}} = 0 \quad \Longrightarrow \quad \ddot{x} + \omega^2 x = 0$$

Momentum and Hamilton function:

$$p = \frac{\partial L}{\partial \dot{x}} = m\dot{x} \quad H = p\dot{x} - L = \frac{p^2}{2m} + \frac{1}{2}m\omega^2 x^2$$

Canonical (Poisson) bracket:

$$\{p, x\} = 1$$

$$\{AB\} = \frac{\partial A}{\partial p} \frac{\partial B}{\partial x} - \frac{\partial A}{\partial x} \frac{\partial B}{\partial p}$$

Canonical quantisation of harmonic oscillator

Canonically conjugated coordinate and momentum become operators:

$$p \rightarrow \hat{p} \quad x \rightarrow \hat{x}$$

Canonical (Poisson) bracket turns to quantum commutator:

$$\{p, x\} = 1 \quad \Longrightarrow \quad [\hat{p}, \hat{x}] = \hat{p}\hat{x} - \hat{x}\hat{p} = -i$$

Hamilton function turns to operator in Hilbert space of wave functions:

$$\hat{H}\psi_n(x) = E_n\psi_n(x)$$

Spectrum of harmonic oscillator:

$$E_n = \omega \left(n + \frac{1}{2} \right) \quad n = 0, 1, 2, \dots$$

Eigenfunctions of harmonic oscillator:

$$\psi_n(x) = \frac{1}{\sqrt{2^n n!}} \left(\frac{m\omega}{\pi} \right)^{1/4} \exp\left(-\frac{m\omega x^2}{2}\right) H_n(\sqrt{m\omega}x)$$

Harmonic oscillator in Holomorphic representation

Equation of motion:

$$\ddot{x} + \omega^2 x = 0$$

General solution (coordinate is real $\implies x^* = x$) and canonical momentum:

$$x(t) = \frac{1}{\sqrt{2m\omega}}(ae^{-i\omega t} + a^*e^{i\omega t}) \equiv \frac{1}{\sqrt{2m\omega}}(a(t) + a^*(t))$$

$$p(t) = m\dot{x}(t) = -i\sqrt{\frac{m\omega}{2}}(ae^{-i\omega t} - a^*e^{i\omega t}) = -i\sqrt{\frac{m\omega}{2}}(a(t) - a^*(t))$$

Proceed from x and p to a and a^* :

$$a(t) = \sqrt{\frac{m\omega}{2}}x + i\frac{p}{\sqrt{2m\omega}} \quad a^*(t) = \sqrt{\frac{m\omega}{2}}x - i\frac{p}{\sqrt{2m\omega}}$$

$$H = \omega \left[\zeta a^* a + (1 - \zeta) a a^* \right] \text{ with any } 0 \leq \zeta \leq 1$$

Quantisation:

$$\{a, a^*\} = 1 \implies [aa^\dagger] = 1$$

Then only $\zeta = \frac{1}{2}$ gives the quantum spectrum consistent with experiment

What is constrained system?

How to restrict the motion of a 3D harmonic oscillator with the Hamiltonian

$$H = \frac{\mathbf{p}^2}{2m} + \frac{1}{2}m\omega^2\mathbf{r}^2$$

to move in the 2D plane (x, y) ?

Way 1: Solve the equations of motion

$$\ddot{\mathbf{r}} + \omega^2\mathbf{r} = 0 \quad \implies \quad \mathbf{r} = \mathbf{A} \sin(\omega t) + \mathbf{B} \cos(\omega t)$$

and then set $A_z = B_z = 0$

Way 2: Impose a constraint $\varphi_1 = z \approx 0$ on the system and ensure its conservation:

$$\varphi_2 = \dot{\varphi}_1 = \frac{\partial \varphi_1}{\partial t} + \{\varphi_1, H\} = \frac{\partial \varphi_1}{\partial \mathbf{p}} \frac{\partial H}{\partial \mathbf{r}} - \frac{\partial \varphi_1}{\partial \mathbf{r}} \frac{\partial H}{\partial \mathbf{p}} = -\frac{p_z}{m} \approx 0$$

Since $\varphi_3 \propto \varphi_1 \approx 0$, the system of constraints is closed and $\{\varphi_1, \varphi_2\} \neq 0$

\implies use φ_1 and φ_2 to get rid of unphysical variables z and p_z :

$$H = \frac{1}{2m}(p_x^2 + p_y^2) + \frac{1}{2}m\omega^2(x^2 + y^2)$$

Free relativistic particle: Different forms of dynamics

Consider a free relativistic particle described by the action

$$S = \int_{\tau_i}^{\tau_f} d\tau \mathcal{L} \quad \mathcal{L} = -m\sqrt{\dot{x}^2} \quad \dot{x}_\mu = \frac{\partial x_\mu}{\partial \tau}$$

Let us proceed from Lagrangian to Hamiltonian:

$$p_\mu = \frac{\partial \mathcal{L}}{\partial \dot{x}^\mu} = -m \frac{\dot{x}_\mu}{\sqrt{\dot{x}^2}} \implies (p\dot{x}) = \mathcal{L} \implies \mathcal{H} = (p\dot{x}) - \mathcal{L} = 0$$

On the other hand, dynamics of the particle is constrained by mass shell condition

$$\varphi = p^2 - m^2 \approx 0$$

Quantisation of the system (Klein–Gordon equation for scalar particle):

$$(\square + m^2) \Psi = 0$$

Reparametrisation invariance: the action is invariant under time redefinition

$$\tau \rightarrow f(\tau) \quad \dot{f}(\tau) > 0 \quad f(\tau_i) = \tau_i \quad f(\tau_f) = \tau_f$$

Alternative approach to quantisation — fix the freedom by choosing time τ :

$$\tau = x_0 \implies \mathcal{L} = -m\sqrt{1 - \dot{\mathbf{r}}^2} \implies \mathcal{H} = \sqrt{\mathbf{p}^2 + m^2}$$

Some definitions and ideas (a la Dirac)

- Constraints **imposed** on the system are called **primary** constraints
- Constraints that guarantee **conservation in time** of the primary constraints are called **secondary** constraints
- Constraints that **commute** with each other (their Poisson bracket vanishes) are called constraints of the **first class**
- The matrix $C_{ab} = \{\varphi_a, \varphi_b\}$ composed of the Poisson brackets of constraints with each other is called **Hessian** matrix
- Constraints such that $\det(C) \neq 0$ are called constraints of the **second class**
- Constraints of the **second class** can be worked out by removing **redundant variables** from the system
- Primary constraints of the **first class** imply **symmetry/invariance** of the system
- Constraints of the **first class** are imposed as **weak (operator)** equations for the wave function after quantisation
- Alternatively, **additional** primary constraints can be imposed on the system to **fix** constraints of the first class to become constraints of the second class (**gauge fixing**)

Electromagnetic field as constrained system

Canonical momentum of electromagnetic field ($\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^2$)

$$\Pi^\mu(x) = \frac{\partial \mathcal{L}}{\partial \dot{A}_\mu(x)} = F^{\mu 0}(x) \quad \Longrightarrow \quad \varphi_1 = \Pi^0 \approx 0 \quad \mathbf{\Pi}(x) = -\dot{\mathbf{A}} - \nabla A_0 = \mathbf{E}$$

Hamiltonian takes the form ($\mathbf{H} = [\nabla \times \mathbf{A}]$)

$$\mathcal{H} = \Pi^\mu \dot{A}_\mu - \mathcal{L} = \int d^3x \left(\frac{1}{2}(\mathbf{E}^2 + \mathbf{H}^2) + A_0(\rho - \nabla \cdot \mathbf{E}) - \mathbf{j} \cdot \mathbf{A} \right)$$

Hamiltonian equations of motion

$$\varphi_2 = \dot{\Pi}_0(x) = \{\Pi_0, \mathcal{H}\} = -\frac{\delta \mathcal{H}}{\delta A_0} \approx 0 \quad \Longrightarrow \quad \nabla \cdot \mathbf{E} = \rho \quad (\text{Gauss law})$$

$$\dot{\mathbf{A}} = \{\mathbf{A}, \mathcal{H}\} = -\frac{\delta \mathcal{H}}{\delta \mathbf{E}} = -\mathbf{E} - \nabla A_0 \approx 0 \quad \Longrightarrow \quad \mathbf{E} = -\dot{\mathbf{A}} - \nabla A_0$$

$$\dot{\mathbf{\Pi}} = \{\mathbf{\Pi}, \mathcal{H}\} = \frac{\delta \mathcal{H}}{\delta \mathbf{A}} = [\nabla \times \mathbf{H}] - \mathbf{j} \approx 0 \quad \Longrightarrow \quad \text{rot} \mathbf{H} - \dot{\mathbf{E}} = \mathbf{j}$$

Gauge invariance

$$\{\varphi_1, \varphi_2\} \approx 0$$

Free relativistic particle: Einbein field formalism

Modify the Lagrange function and introduce einbein field μ (with **no** velocity $\dot{\mu}$)

$$\mathcal{L} = -\frac{m^2}{2\mu} - \frac{\mu\dot{x}^2}{2} \quad \xRightarrow{\mu \rightarrow \mu_{\text{ext}}} \quad \mathcal{L} = -m\sqrt{\dot{x}^2}$$

Evaluate canonical momenta for coordinates x_μ and μ

$$p_\mu = \frac{\partial L}{\partial \dot{x}^\mu} = -\mu \dot{x}_\mu \quad \pi = \frac{\partial L}{\partial \dot{\mu}} = 0 \quad \Longrightarrow \quad \varphi_1 = \pi \approx 0$$

The Hamiltonian takes the form

$$\mathcal{H} = -\frac{1}{2\mu}(p^2 - m^2) + \lambda\varphi_1 \quad \Longrightarrow \quad \varphi_2 = \{\varphi_1, \mathcal{H}\} = \frac{1}{2\mu^2}(p^2 - m^2) \approx 0$$

Quantisation follows the standard lines

$$[\hat{p}_\mu \hat{x}_\nu] = -ig_{\mu\nu} \quad \hat{p}_\mu = -i\frac{\partial}{\partial x^\mu} \quad [\hat{\pi} \hat{\mu}] = -1 \quad \hat{\pi} = \frac{\partial}{\partial \mu}$$

To arrive at

$$(\square + m^2)\Psi = 0 \quad \frac{\partial \Psi}{\partial \mu} = 0$$

Free relativistic particle: Gauge fixing

To fix laboratory gauge impose an extra constraint

$$\varphi_3 = x_0 + \tau \quad \xRightarrow{x_0 \rightarrow x'_0 = x_0 + \tau} \quad \varphi_3 = x'_0$$

The partition function for this canonical transformation takes the form

$$F(x, p', \tau) = p'_\mu (x^\mu + \tau g^{\mu 0})$$

and the modified Hamilton function becomes

$$\mathcal{H}' = \mathcal{H} + \frac{\partial F}{\partial \tau} = \mathcal{H} + p'_0$$

On the constraints surface

$$p'^2 \approx m^2 \quad \Longrightarrow \quad p'_0 = \sqrt{\mathbf{p}^2 + m^2} \quad \& \quad \mathcal{H} \approx 0$$

Therefore, after fixing the gauge of the laboratory time we have

$$\mathcal{H}' = \sqrt{\mathbf{p}^2 + m^2}$$

Two noninteracting relativistic particles

Start from the Lagrange function and introduce two einbeins

$$\mathcal{L} = -m_1 \sqrt{\dot{x}_1^2} - m_2 \sqrt{\dot{x}_2^2} \implies -\frac{\mu_1 \dot{x}_1^2}{2} - \frac{m_1^2}{2\mu_1^2} - \frac{\mu_2 \dot{x}_2^2}{2} - \frac{m_2^2}{2\mu_2^2}$$

Separate the centre-of-mass motion (formal **nonrelativistic** relations apply!)

$$x_\mu = x_{1\mu} - x_{2\mu} \quad X_\mu = \zeta x_{1\mu} + (1 - \zeta)x_{2\mu} \quad \zeta = \frac{\mu_1}{\mu_1 + \mu_2} \quad M = \mu_1 + \mu_2$$

In terms of the new variables the Lagrange function becomes

$$\mathcal{L} = -\frac{m_1^2}{2M\zeta} - \frac{m_2^2}{2M(1-\zeta)} - \frac{1}{2}M \left(\dot{X} - \dot{\zeta}x \right)^2 - \frac{1}{2}M\zeta(1-\zeta)\dot{x}^2$$

Then we define the four canonically conjugated momenta

$$P_\mu = \frac{\partial L}{\partial \dot{X}_\mu} \quad p_\mu = \frac{\partial L}{\partial \dot{x}_\mu} \quad \Pi = \frac{\partial L}{\partial \dot{M}} \quad \pi = \frac{\partial L}{\partial \dot{\zeta}}$$

and proceed to the Hamilton function

$$\mathcal{H} = \frac{m_1^2 - p^2}{2M\zeta} + \frac{m_2^2 - p^2}{2M(1-\zeta)} - \frac{P^2}{2M} + \Lambda\varphi_1 + \lambda\varphi_2 \quad \varphi_1 = \Pi \quad \varphi_2 = \pi + (Px)$$

Two noninteracting relativistic particles

The Poisson bracket is defined as

$$\{AB\} = \frac{\partial A}{\partial P_\mu} \frac{\partial B}{\partial X_\mu} - \frac{\partial A}{\partial X_\mu} \frac{\partial B}{\partial P_\mu} + \frac{\partial A}{\partial p_\mu} \frac{\partial B}{\partial x_\mu} - \frac{\partial A}{\partial x_\mu} \frac{\partial B}{\partial p_\mu} + \frac{\partial A}{\partial \Pi} \frac{\partial B}{\partial M} - \frac{\partial A}{\partial M} \frac{\partial B}{\partial \Pi} + \frac{\partial A}{\partial \pi} \frac{\partial B}{\partial \mu} - \frac{\partial A}{\partial \mu} \frac{\partial B}{\partial \pi}$$

and the secondary constraints take the form

$$\varphi_3 = \{\varphi_1 H\} = -\frac{\varepsilon_1^2}{2M^2\zeta} - \frac{\varepsilon_2^2}{2M^2(1-\zeta)} + \frac{P^2}{2M^2}$$

$$\varphi_4 = \{\varphi_2 H\} = -\frac{\varepsilon_1^2}{2M\zeta^2} + \frac{\varepsilon_2^2}{2M(1-\zeta)^2} + \frac{(pP)}{M\zeta(1-\zeta)}$$

$$\varepsilon_1^2 = m_1^2 - p^2 \quad \varepsilon_2^2 = m_2^2 - p^2$$

Constraints $\{\varphi_1, \varphi_2, \varphi_3, \varphi_4\}$ are **first class** constraints (2 primary and 2 secondary)

\implies There are **two "gauge" freedoms**: for c.m. motion & for relative motion

Time-like gauge

Two fix the two gauges, impose two **extra** constraints

$$\varphi_5 = X_0 + \tau \quad \varphi_6 = x_0$$

and “**commute**” them with the Hamilton function for secondary constraints

$$\varphi_7 = \{\varphi_5 H\} = \frac{P_0}{M} - 1 \quad \varphi_8 = \{\varphi_6 H\} = \frac{p_0}{M\zeta(1-\zeta)}$$

The Hamilton function becomes simply

$$H = M = P_0$$

and the canonical bracket has to be redefined: **Poisson** bracket \implies **Dirac** bracket:

$$\{AB\}^* = \{AB\} - \sum_{a,b} \{A\varphi_a\} C_{ab}^{-1} \{\varphi_b B\}$$

where C^{-1} is the inverse matrix with respect to $(a, b = 1..8)$

$$C_{ab} = \{\varphi_a \varphi_b\}$$

Time-like gauge

The Hamiltonian takes the form

$$\mathcal{H} = \sqrt{\mathbf{P}^2 + \mathcal{E}_1^2 + \mathcal{E}_2^2 + 2\sqrt{\mathcal{E}_1^2 \mathcal{E}_2^2 + (\mathbf{p}\mathbf{P})^2}}$$

where $\mathcal{E}_i = \sqrt{\mathbf{p}^2 + m_i^2}$ ($i = 1, 2$) and the set of canonical (Dirac) brackets reads

$$\{P_i X_k\}^* = \delta_{ik}$$

$$\{X_i p_k\}^* = \frac{p_i P_k}{M^2}$$

$$\{X_i X_k\}^* = \frac{x_i p_k - x_k p_i}{M^2}$$

$$\{X_i x_k\}^* = \frac{x_i P_k}{M^2} + \frac{x_i p_k}{M^2} \left(\frac{1-\zeta}{\zeta} - \frac{\zeta}{1-\zeta} \right)$$

$$\{p_i x_k\}^* = \delta_{ik} - \frac{P_i P_k}{M^2} - \frac{P_i p_k}{M^2} \left(\frac{1-\zeta}{\zeta} - \frac{\zeta}{1-\zeta} \right)$$

$$\{P_i P_k\}^* = \{P_i p_k\}^* = \{P_i x_k\}^* = \{p_i p_k\}^* = \{x_i x_k\}^* = 0$$

Newton–Wigner coordinate

Let us define the Newton–Wigner coordinate

$$Q_i = X_i - \frac{S_{ik}P_k}{E(M+E)} \quad E = \sqrt{M^2 - P^2} \quad S_{ik} = x_i p_k - x_k p_i$$

and the proper internal variables

$$\mathbf{k} = \mathbf{p} + \frac{(\mathbf{p}\mathbf{P})\mathbf{P}}{E(M+E)} \quad \mathbf{r} = \mathbf{x} + \frac{(\mathbf{x}\mathbf{P})\mathbf{P}}{E(M+E)} + \frac{(\mathbf{x}\mathbf{P})\mathbf{k}}{E\omega_1\omega_2} \left(\omega_1 - \omega_2 - \frac{(\mathbf{k}\mathbf{P})}{M} \right)$$

where

$$M = \sqrt{P^2 + (\omega_1 + \omega_2)^2} \quad \omega_1 = \sqrt{m_1^2 + \mathbf{k}^2} \quad \omega_2 = \sqrt{m_2^2 + \mathbf{k}^2}$$

with the canonical Dirac brackets:

$$\{k_i r_k\}^* = \delta_{ik} \quad \{k_i k_k\}^* = \{r_i r_k\}^* = \{k_i Q_k\}^* = \{r_i Q_k\}^* = 0$$

and the Hamilton function takes the desired form

$$\mathcal{H} = M = \sqrt{P^2 + \left(\sqrt{m_1^2 + \mathbf{k}^2} + \sqrt{m_2^2 + \mathbf{k}^2} \right)^2}$$

Proper-time gauge

Let us fix one gauge freedom of two imposing Lorenz-covariant constraint φ_5 that gives rise to a covariant secondary constraint φ_6 :

$$\varphi_5 = (Px) \approx 0 \quad \implies \quad \varphi_6 = (Pp) \approx 0$$

After excluding 6 second-class constraints, Newton–Wigner coordinate is defined as

$$Z_\mu = X_\mu + \frac{1}{2} S_{ij} \Gamma_{ij\mu} \quad \{Z_\mu Z_\nu\}^* = 0 \quad \{P_\mu Z_\nu\}^* = g_{\mu\nu}$$

with $S_{ij} = e_{i\mu} e_{j\nu} S_{\mu\nu}$ and a tetrade of vectors and Christoffel symbols introduced as

$$e_{0\mu} = \frac{P_\mu}{\sqrt{P^2}} \quad e_{i\mu} e_{j\mu} = -\delta_{ij} \quad e_{0\mu} e_{j\mu} = 0 \quad \Gamma_{ij\alpha} = e_{i\mu} \frac{\partial}{\partial P_\alpha} e_{j\mu} \quad \Gamma_{0j\alpha} = e_{0\mu} \frac{\partial}{\partial P_\alpha} e_{j\mu}$$

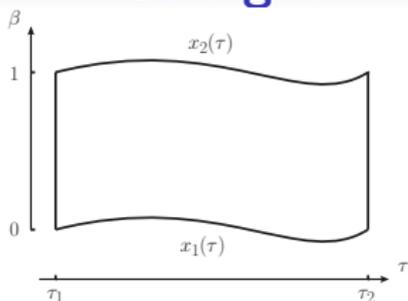
Then the coordinate and momentum of relative motion can be defined as

$$x_i = -e_{i\mu} x^\mu \quad p_i = -e_{i\mu} p^\mu$$

and the system is quantised employing the remaining primary first-class constraint

$$\left(\hat{P}^2 - \mathcal{M}^2(\hat{p}_i) \right) \Psi = 0 \quad \mathcal{M}^2 = \left(\sqrt{m_1^2 + p_i^2} + \sqrt{m_2^2 + p_i^2} \right)^2$$

Straight-line string: Lagrange function



$$S_{\text{string}} = -\sigma \int_{\tau_1}^{\tau_2} d\tau \int_0^1 d\beta \sqrt{(\dot{w}w')^2 - \dot{w}^2 w'^2}$$

$$w_\mu(\beta, \tau) = (1 - \beta)x_{1\mu}(\tau) + \beta x_{2\mu}(\tau)$$

$$\dot{w}_\mu = \frac{\partial w_\mu}{\partial \tau} \quad w'_\mu = \frac{\partial w_\mu}{\partial \beta} = x_{2\mu} - x_{1\mu} \equiv x_\mu$$

Continuous einbein field $\nu(\beta)$ is introduced to get rid of the square root

$$L = -\frac{1}{2} \int_0^1 d\beta \nu(\beta) \left(\dot{w}^2 - \frac{(\dot{w}x)^2}{x^2} \right) + \frac{1}{2} \int_0^1 d\beta \frac{\sigma^2 x^2}{\nu(\beta)}$$

with canonical momenta ($X_\mu = \zeta x_{1\mu} + (1 - \zeta)x_{2\mu}$) with $\zeta = \int_0^1 d\beta \nu \beta / \int_0^1 d\beta \nu$)

$$P_\mu = \frac{\partial L}{\partial \dot{X}^\mu} = -M \left(\dot{X}_\mu - \frac{x_\mu (x \dot{X})}{x^2} \right) \implies \varphi_1 = (Px) \approx 0$$

$$p_\mu = \frac{\partial L}{\partial \dot{x}^\mu} = -m \left(\dot{x}_\mu - \frac{x_\mu (x \dot{x})}{x^2} \right) \implies \varphi_2 = (px) \approx 0$$

$$\kappa(\beta) = \frac{\delta L}{\delta \dot{\nu}(\beta)} = 0 \implies \varphi_3(\beta) = \kappa(\beta) \approx 0$$

Straight-line string: Hamilton function

The Hamilton function takes the form

$$\mathcal{H} = \underbrace{-\frac{P^2}{2M} - \frac{p^2}{2m} - \frac{kx^2}{2}}_{\mathcal{H}_0} + \underbrace{\Lambda(Px)}_{\varphi_1} + \underbrace{\lambda(px)}_{\varphi_2} + \int_0^1 d\beta e(\beta) \underbrace{\kappa(\beta)}_{\varphi_3(\beta)}$$

$$M = \int_0^1 d\beta \nu \quad m = \int_0^1 d\beta \nu (\beta - \zeta)^2 \quad k = \int_0^1 \frac{\sigma^2}{\nu}$$

The Poisson bracket is defined as

$$\{AB\} = \frac{\partial A}{\partial P_\mu} \frac{\partial B}{\partial X_\mu} - \frac{\partial A}{\partial X_\mu} \frac{\partial B}{\partial P_\mu} + \frac{\partial A}{\partial p_\mu} \frac{\partial B}{\partial x_\mu} - \frac{\partial A}{\partial x_\mu} \frac{\partial B}{\partial p_\mu} + \int_0^1 d\beta \left(\frac{\delta A}{\delta \kappa(\beta)} \frac{\delta B}{\delta \nu(\beta)} - \frac{\delta A}{\delta \nu(\beta)} \frac{\delta B}{\delta \kappa(\beta)} \right)$$

Primary constraints lead to secondary ones and finally Lagrange multipliers get fixed

$$\left. \begin{aligned} \varphi_1 \approx 0 &\implies \varphi_4 = \{\varphi_1, \mathcal{H}\} \approx 0 \implies \{\varphi_4, \mathcal{H}\} \approx 0 \\ \varphi_2 \approx 0 &\implies \varphi_5 = \{\varphi_2, \mathcal{H}\} \approx 0 \implies \{\varphi_5, \mathcal{H}\} \approx 0 \\ \varphi_3 \approx 0 &\implies \varphi_6 = \{\varphi_3, \mathcal{H}\} \approx 0 \implies \{\varphi_6, \mathcal{H}\} \approx 0 \end{aligned} \right\} \implies \begin{cases} \Lambda = 0 \\ \lambda = 0 \\ e(\beta) = e_0 \nu(\beta) \end{cases}$$

The remaining pair of first-class constraints is

$$\Phi_1 = \int_0^1 d\beta \nu(\beta) \varphi_3(\beta) \quad \Phi_2 = \int_0^1 d\beta \nu(\beta) \varphi_6(\beta) \approx -\mathcal{H}_0$$

Straight-line string: Quantisation and spectrum

We introduce the standard Newton–Wigner coordinate

$$Z_\mu = X_\mu + \frac{1}{2}S_{ij}\Gamma_{ij\mu}$$

where

$$n_i = -e_{i\mu} \frac{x_\mu}{\sqrt{-x^2}} \quad S_{ij} = e_{i\mu}e_{j\nu}S_{\mu\nu} \quad S_{\mu\nu} = x_\mu p_\nu - x_\nu p_\mu \quad L^2 = \frac{1}{2}S_{ik}S_{ik}$$

and tetrade of vectors and Christoffel symbols are defined as before

$$e_{0\mu} = \frac{P_\mu}{\sqrt{P^2}} \quad e_{i\mu}e_{j\mu} = -\delta_{ij} \quad e_{0\mu}e_{j\mu} = 0 \quad \Gamma_{ij\alpha} = e_{i\mu} \frac{\partial}{\partial P_\alpha} e_{j\mu} \quad \Gamma_{0j\alpha} = e_{0\mu} \frac{\partial}{\partial P_\alpha} e_{j\mu}$$

Employing second-class constraints to exclude redundant degrees of freedom we find

$$\nu(\beta) = \frac{N}{\sqrt{1 - (2\beta - 1)^2}} \quad M = 8m = \frac{\pi N}{2} \quad k = \frac{\pi\sigma^2}{4N} \quad x^2 = -\frac{8L}{\pi\sigma} \quad p^2 = -\frac{\pi\sigma L}{8}$$

with arbitrary N . After quantisation we get for bound state equation and spectrum

$$(\hat{P}^2 - 2\pi\sigma\hat{L})\Psi = 0 \quad \implies \quad \mathcal{M}_L^2 = 2\pi\sigma\sqrt{L(L+1)}$$

Generalisation to fermion in external field

Spin- $\frac{1}{2}$ fermion in external electromagnetic field is described by Lagrange function

$$\mathcal{L} = -\frac{\mu}{2}\dot{x}_\nu[\dot{x}^\nu - i\chi\psi^\nu] - \frac{i}{2}\psi_\nu\dot{\psi}^\nu - \frac{m^2}{2\mu} + \frac{i}{2}[\psi_5\dot{\psi}_5 + m\chi\psi_5] - eA_\mu\dot{x}^\mu + \frac{ie}{2\mu}\psi_\mu\psi_\nu F^{\mu\nu}$$

where ψ_μ describes the spin and μ and χ are einbeins (ψ_μ and χ are **Grassmannian**)

Balachandran, Salomonson, Skagerstam, Winnberg'1977

Two gauge-like symmetries of the action:

- Reparametrisation invariance

$$\tau \rightarrow f(\tau) \quad \mu \rightarrow \frac{\mu}{\dot{f}(\tau)} \quad \chi \rightarrow \frac{\chi}{\dot{f}(\tau)}$$

- Supergauge transformations (α is Grassmannian)

$$\delta x_\nu = i\alpha\psi_\nu \quad \delta\psi_\nu = -\alpha\mu(\dot{x}_\nu - \frac{i}{2}\chi\psi_\nu) \quad \delta\mu = i\alpha\mu^2\chi \quad \delta\chi = 2\dot{\alpha} \quad \delta\psi_5 = m\alpha$$

Gitman, Tyutin'1990; Grigoryan, Grigoryan'1991

Quantisation with first-class constraints

Quantisation proceeds via setting

$$\hat{p}_\mu = i \frac{\partial}{\partial x_\mu} \quad \psi_\mu = \frac{1}{\sqrt{2}} \gamma_5 \gamma_\mu \quad \psi_5 = \frac{1}{\sqrt{2}} \gamma_5$$

Primary constraints $\Phi_1 = \pi \approx 0$ and $\Phi_2 = \pi_\chi \approx 0$ give rise to secondary **first-class** constraints Φ_{KG} and Φ_D that, after quantisation, transform into equations

$$\hat{\Phi}_D \Psi = \gamma_5 \left(\gamma(p + eA) - m \right) \Psi = 0$$

$$\hat{\Phi}_{KG} \Psi = \left((p + eA)^2 - m^2 - \frac{ie}{2} \sigma_{\mu\nu} F_{\mu\nu} \right) \Psi = 0$$

These are the famous Dirac equation and so-called “Klein–Gordon with spin”

Quantum algebra of constraints is closed:

$$[\hat{\Phi}_{KG}, \hat{\Phi}_{KG}]_- = 0 \quad [\hat{\Phi}_{KG}, \hat{\Phi}_D]_- = 0 \quad [\hat{\Phi}_D, \hat{\Phi}_D]_+ = -\hat{\Phi}_{KG}$$

Quantisation after “laboratory gauge” fixing

Fix “gauge” freedom imposing additional primary constraint

$$\tau = x_0$$

Then, after some algebra, the Hamilton function and Dirac constraint take the form

$$H = \mu_0 + eA_0 + \frac{ie}{\mu_0} F_{0i} \psi_0 \psi_i \quad \varphi_D = \mu_0 \psi_0 - (\mathbf{p} - e\mathbf{A})\psi + m\psi_5$$

where the einbein field takes the form (solution with $\mu_0 < 0$ thrown away!)

$$\mu_0 = \sqrt{(\mathbf{p} - e\mathbf{A})^2 + m^2 - ieF_{ik}\psi_i\psi_k}$$

However, the quantum algebra is not closed:

$$[\hat{H}, \hat{\varphi}_D]_- \neq 0$$

since the Grassmann algebra for ψ 's

$$\psi_\mu \psi_\nu = -\psi_\nu \psi_\mu \quad \psi_5 \psi_\mu = -\psi_\mu \psi_5 \quad \psi_5 \psi_5 = 0$$

does **not** translate into similar algebra for Dirac matrices

The physical origins of the problem

Start from Dirac equation in external field (for simplicity, $\mathbf{A} = 0$) and perform its Foldy–Wouthuysen transformation to order $1/m^2$

$$H = \frac{p^2}{2m} - \frac{p^4}{8m^3} + eA_0 - \underbrace{\frac{e}{4m^2} \boldsymbol{\sigma} [\mathbf{E} \mathbf{p}]}_{\text{Spin-orbit term}} - \underbrace{\frac{e}{8m^2} \text{div} \mathbf{E}}_{\text{Darwin term}}$$

where

$$V_{\text{Darwin}} = -\frac{e}{8m^2} \text{div} \mathbf{E} \propto \Delta A_0 \propto \rho_{\text{ext}}(\mathbf{x})$$

and, in central field $A_0(r)$,

$$\mathbf{E} = -\frac{r}{r} \frac{dA_0}{dr}$$

- Quantisation with **first class** constraints \implies Darwin term is **present**
- Quantisation after **laboratory time** fixing \implies Darwin is **missing**

Conclusion: **Zitterbewegung** (motion with $\mu_0 < 0$) lost after fixing **laboratory time**
 \implies **no problem** for **free** particles but potential **danger** for **interacting** particles

Summary: How to quantise constrained system

- Start from **Lagrange** function
- Evaluate canonical **momenta**
- Establish **primary** constraints
- Build **Hamilton** function
- Add **primary** constraints to **Hamilton** function with Lagrange **multiplies**
- **Commute** **primary** constraints with Hamilton function for **secondary** constraints
- If the process **truncates** at some step **proceed** to the next stage (give up if not)
- **Commute** (build Poisson brackets) **all** constraints with each other
- **Identify** **first-** and **second-class** constraints
- **Choose** whether or not to **fix** the first-class constraints
- **Impose** additional constraints to fix the chosen first-class constraints (**Caution!**)
- **Establish** the corresponding **secondary** constraints
- **Get rid** of **second-class** constraints by eliminating unphysical degrees of freedom
- Build **Hessian matrix** and **Dirac brackets** for all **physical** degrees of freedom
- **Bring** Dirac brackets to **canonical** form by defining the **proper** coordinates
- **Express** everything through the **physical** variables
- **Quantise** system employing **Hamilton** function and remaining **first-class** constraints
- **Check** if the resulting theory describes Nature — if yes, **enjoy** your theory!

Conclusions

- Quantisation of a classical system is an **Art**
 - Any **quantum theory** possesses the **only** classical limit
 - The same **classical theory** may correspond to an **infinite number** of quantum theories (operators ordering, gauge fixing, etc)
- Presence of **constraints** makes quantisation an **Art²**
- **General principles** for quantisation of constrained systems can still be formulated that work in practical **applications** and lead to **meaningful** theories

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Thank you for your attention!

Possible choices for tetrads

Orthogonality and completeness conditions

$$e_{\mu\nu}e_{\mu'\nu} = g_{\mu\mu'} \quad e_{\mu\nu}e_{\mu\nu'} = g_{\nu\nu'}$$

where μ is serial number of tetrad in the set and ν is its Lorentz component

- Simplest choice

$$e_{\mu\nu} = g_{\mu\nu}$$

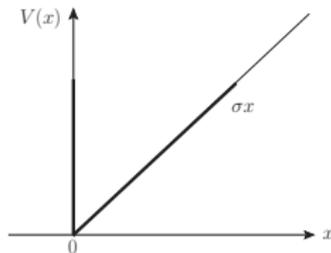
- Convenient choice for co-moving frame

$$e_{0\mu} = \frac{P_\mu}{\sqrt{P^2}} \quad e_{i0} = \frac{P_i}{\sqrt{P^2}} \quad e_{ik} = \delta_{ik} + \frac{P_i P_k}{\sqrt{P^2}(P_0 + \sqrt{P^2})}$$

Einbein field as variational parameter

One-dimensional linear potential

$$V(x) = \begin{cases} \sigma x, & x \geq 0 \\ \infty, & x < 0 \end{cases}$$



Schrödinger equation and wave function

$$\left(\frac{p^2}{2m} + V(x) \right) \psi = E\psi \quad \psi_n(x) = \mathcal{N}_n Ai(y) \quad y = (2m\sigma)^{1/3} (x - E_n/\sigma)$$

Quantisation condition

$$\psi_n(x=0) = 0 \quad \Longrightarrow \quad E_n = -\frac{\sigma^{2/3}}{(2m)^{1/3}} a_n \quad \text{with} \quad Ai(a_n) = 0$$

Approximate formula for zeros of Airy function

$$a_n \approx - \underbrace{\left(\frac{3\pi}{2} \right)^{2/3}}_{\approx 2.81} \left(n - \frac{1}{4} \right)^{2/3} \quad n = 1, 2, \dots$$

Einbein field as variational parameter

$V(x)$ ↑

One

In einbein field formalism:

$$\sigma x \rightarrow \frac{\nu}{2} + \frac{1}{2} m \omega^2 x^2 \quad \omega = \frac{\sigma}{\sqrt{m\nu}}$$

Sch

$\left(\frac{p}{2m}\right)$

$$E_n = \frac{\nu}{2} + \omega \left[\{2(n-1) + 1\} + \frac{1}{2} \right] \Rightarrow -\frac{\sigma^{2/3}}{(2m)^{1/3}} a_n^{\text{ein}} \quad - E_n/\sigma$$

Qua

$$a_n^{\text{ein}} = -3 \left(n - \frac{1}{4} \right)^{2/3} \quad n = 1, 2, \dots$$

$$\psi_n(x=0) = 0 \Rightarrow E_n = -\frac{\sigma^{2/3}}{(2m)^{1/3}} a_n^{\text{ein}} \quad \text{with} \quad A_l(a_n) = 0$$

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