

# Hyperfine Electronic Bridge in the $^{229}\text{Th}$ Nucleus

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2026.01.20

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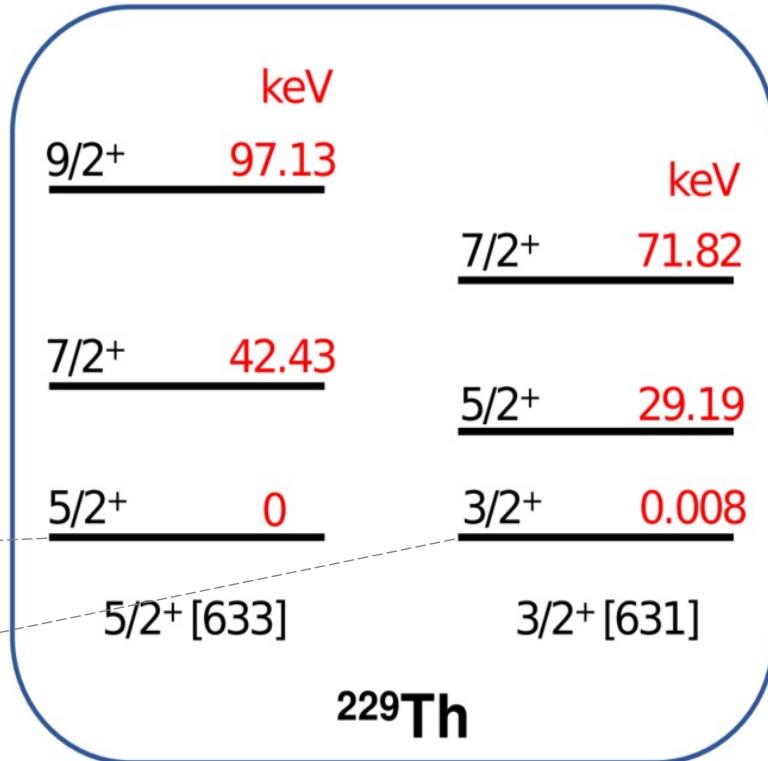
# 1. Introduction to the $^{229}\text{Th}$ isomer

Thorium (Th): atomic number 90

$^{229}\text{Th}$ : a radiative isotope with a half-life  
of about 7900 yr

$^{229\text{m}}\text{Th}$ : a unique isomeric state with  
an energy of about 8 eV

Typical nuclear excitation energies: keV and MeV



Partial nuclear levels of  $^{229}\text{Th}$

W. Wang, H. Zhang, X. Wang, J. Phys. B **54**, 244001 (2021)

# 1. Introduction to the $^{229}\text{Th}$ isomer

Atomic clock: a frequency reference based on a specific transition between atomic levels

Fractional frequency uncertainty:  $10^{-19}$

Limitation: external electromagnetic perturbations

A potential solution:  $^{229}\text{Th}$  nuclear clock

E. Peik, C. E. Tamm, *Europhys. Lett.* **61**, 181 (2003)

- High robustness against external electromagnetic perturbations
- Narrow relative linewidth  $\Delta E/E \sim 1.6 \times 10^{-20}$
- Accessible with a vacuum-ultraviolet (VUV) laser: 8 eV  $\sim 150$  nm

A candidate for next-generation frequency standard

V. V. Flambaum, *Phys. Rev. Lett.* **97**, 092502 (2006)

Precision measurements, search for variations of fundamental constants

# 1. Introduction to the $^{229}\text{Th}$ isomer

Weak radiative transition: lifetime  $\sim 2500$  s

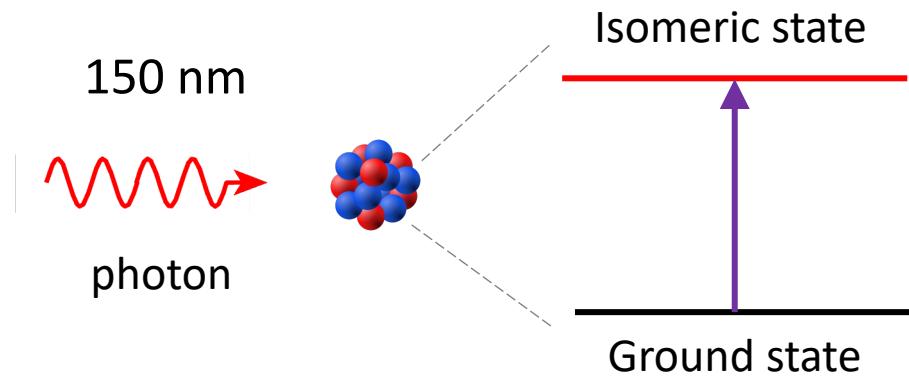
J. Tiedau et al., Phys. Rev. Lett. **132**, 182501 (2024)

It is difficult to excite the  $^{229}\text{Th}$  isomer

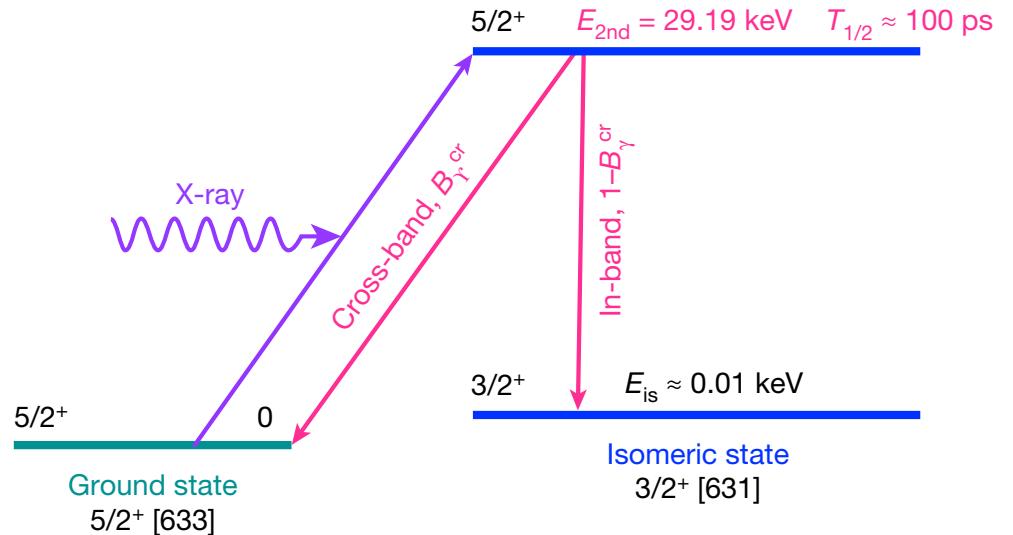
Various methods are proposed

# 1. Introduction to the $^{229}\text{Th}$ isomer

## 1. Direct optical excitation:



## 2. Indirect x-ray pumping:



After many years of effort, it was finally successful.

C. Zhang et al., Nature (London) **633**, 63 (2024)

J. Tiedau et al., Phys. Rev. Lett. **132**, 182501 (2024)

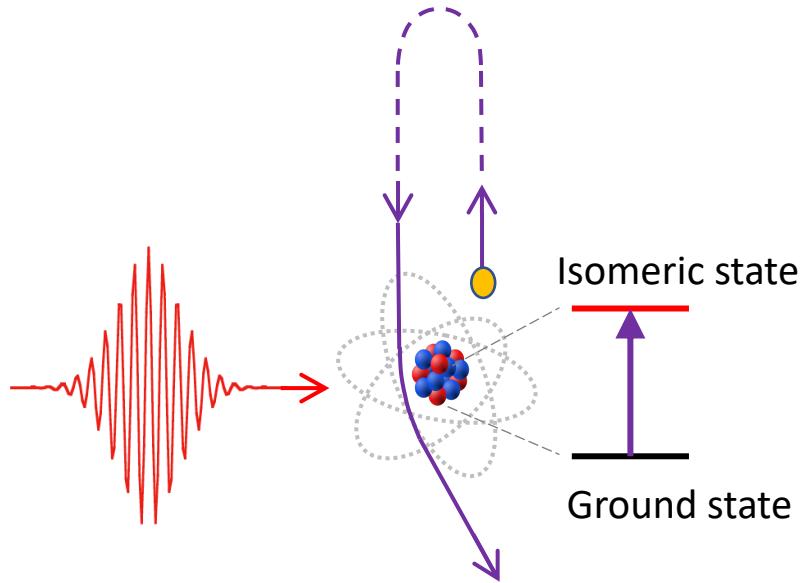
R. Elwell et al., Phys. Rev. Lett. **133**, 013201 (2024)

T. Hiraki et al., Nat. Commun. **15**, 5536 (2024)

T. Masuda et al., Nature (London) **573**, 238 (2019)

# 1. Introduction to the $^{229}\text{Th}$ isomer

## 3. Strong laser field excitation:



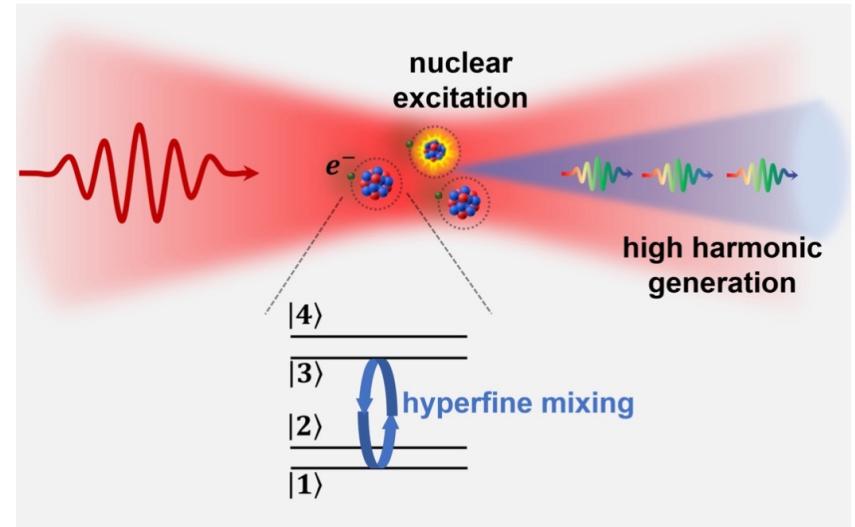
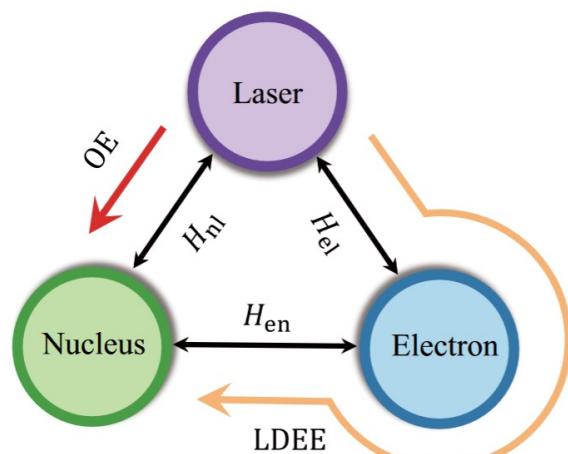
Semi-classical method:  
recollision electron

$\text{Th}^{0+,1+,2+,3+,4+} \sim 10^{-11}$  per pulse

**W. Wang, J. Zhou, B. Liu, and X. Wang,**  
Phys. Rev. Lett. **127**, 052501 (2021)

Quantum theory: laser-nucleus-electron  
interaction

**W. Wang and X. Wang, Phys. Rev.**  
Res. **5**, 043232 (2023)



Nuclear hyperfine mixing

$\text{Th}^{89+} \sim 0.1 - 0.9$  per pulse

**H. Zhang, T. Li, and X. Wang, Phys. Rev. Lett. **133**, 152503 (2024); Phys. Rev. C **111**, 044614 (2025)**

# 1. Introduction to the $^{229}\text{Th}$ isomer

## 4. Electronic bridge

## 5. Other methods

## 2. Electronic bridge

Electronic Bridge (EB):

- Nuclear transition is induced by laser-driven electronic transitions
- Photon energy + Electronic transition energy = Nuclear energy

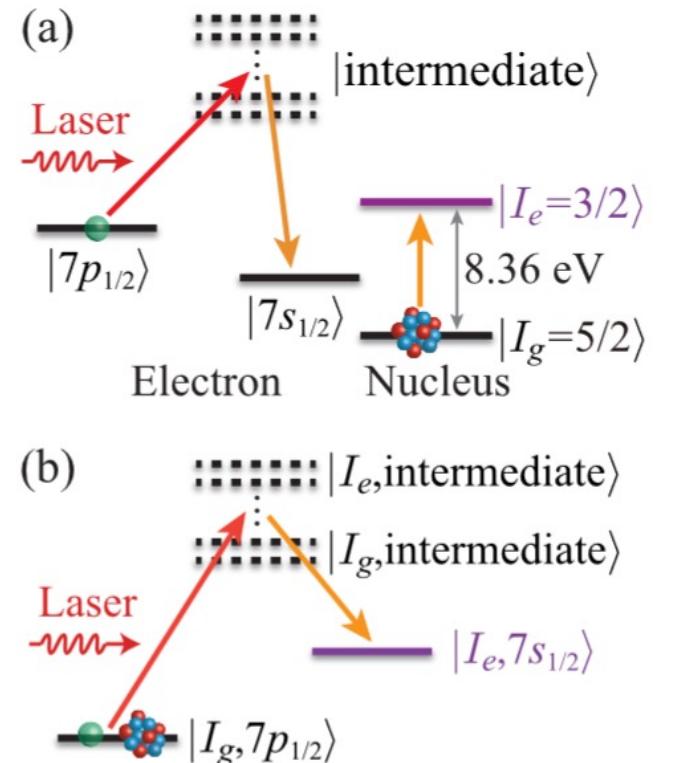
I. S. Batkin, Sov. J. Nucl. Phys. **29**, 464 (1979)

E. V. Tkalya, JETP Lett. **55**, 216 (1992)

- EB can enhance the effective coupling between laser and  $^{229m}\text{Th}$
- Two-photon EB: two lasers, wavelength  $\sim 300$  nm

S. G. Porsev *et al.*, Phys. Rev. Lett. **105**, 182501 (2010)

W. Wang, F. Zou, S. Fritzsche, and Y. Li, Phys. Rev. Lett. **133**, 223001 (2024)



Schematic diagram of EB in  $^{229}\text{Th}^{3+}$

## 2. Electronic bridge

Limitations of existing research:

### 1. Neglecting the hyperfine structure

Neglecting the hyperfine splitting of levels, we can represent the total wave function as a product of the nuclear wave function and the electronic wave function. For instance,

S. G. Porsev and V. V. Flambaum, Phys. Rev. A **81**, 032504 (2010)

Hyperfine splitting in  $^{229}\text{Th}^{3+}$ : GHz

Measurement uncertainty of  $E_n$ : kHz

J. Tiedau *et al.*, Phys. Rev. Lett. **132**, 182501 (2024)

C. Zhang *et al.*, Nature (London) **633**, 63 (2024)

### 2. Excitation rate based on the Fermi's golden rule

Perturbation behavior, incoherent excitation

$$W_{ba}^{\text{in}} = W_{ab} \frac{4\pi^3 c^2}{\omega^3} I_\omega.$$

### 3. For two-photon EB, the decay of excited electronic level is neglected

As a result, the actual efficiency of EB method remains unknown

## 2. Dressed hyperfine state

Atomic level:  $(I, \gamma J)$

$I$  is the nuclear spin,  $J$  is the electronic angular momentum

Hyperfine interaction:  $H_{n-e} = \sum_{\tau K} \mathcal{M}^{(\tau K)} \cdot T^{(\tau K)}, \quad \gamma$  denotes other electronic quantum numbers

With  $H_{n-e}$ , hyperfine level:  $(I, \gamma J, F)$

$F$  is the total angular momentum

Dressed hyperfine state:  $|[I\gamma J]FM\rangle = a|I\gamma J; FM\rangle + \sum_t b_t |I_t \gamma_t J_t; FM\rangle \longleftrightarrow (I, \gamma J, F)$

- Eigenstate of the coupled electron–nucleus system

Hyperfine-coupled basis:  $|I\gamma J; FM\rangle$

- Mixing of different nuclear states, i.e.,

nuclear hyperfine mixing

Mixing coefficient:  $b_t = \sum_{\tau K} \frac{(-1)^{I+J_t+F}}{E_0 - E_t} \begin{Bmatrix} I_t & J_t & F \\ J & I & K \end{Bmatrix} \times \langle I_t | \mathcal{M}^{(\tau K)} | I \rangle \langle \gamma_t J_t | T^{(\tau K)} | \gamma J \rangle,$

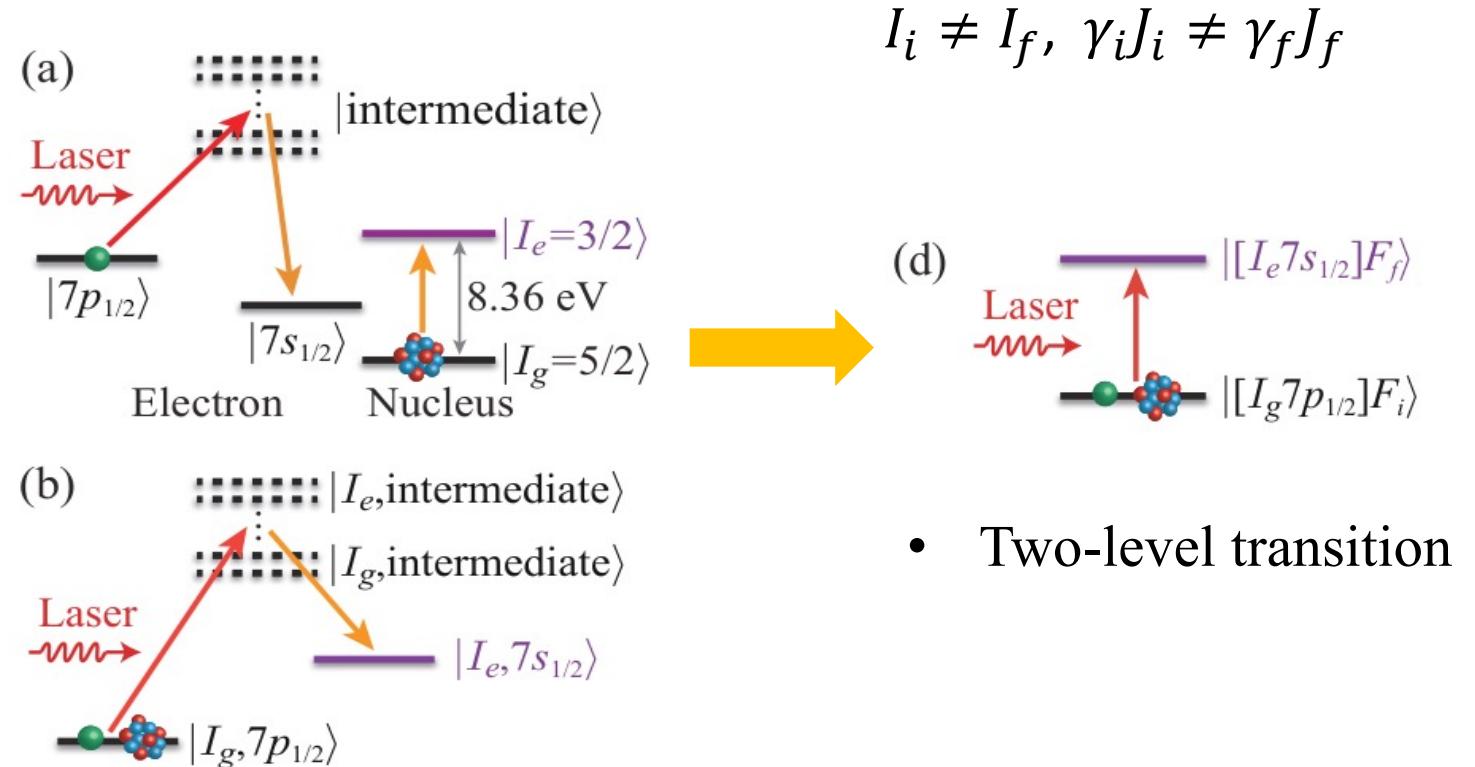
W. Wang, F. Zou, S. Fritzsche, and Y. Li, Phys. Rev. Lett. **133**, 223001 (2024)

## 2. Hyperfine electronic bridge

Hyperfine EB (HEB) transition:  $|[I_i \gamma_i J_i] F_i M_i\rangle \rightarrow |[I_f \gamma_f J_f] F_f M_f\rangle$  or  $(I_i, \gamma_i J_i, F_i) \rightarrow (I_f, \gamma_f J_f, F_f)$

The advantages of HEB:

- Hyperfine structure is considered
- The intermediate state of traditional EB was eliminated
- Easy to combine with the quantum-optical methods



W. Wang, F. Zou, S. Fritzsch, and Y. Li, Phys. Rev. Lett. **133**, 223001 (2024)    W. Wang, S. Fritzsch, and Y. Li, Phys. Rev. A **112**, 022811 (2025)

## 2. Quantum-optical model based on HEB

With the HEB transition, a quantum-optical model is readily constructed

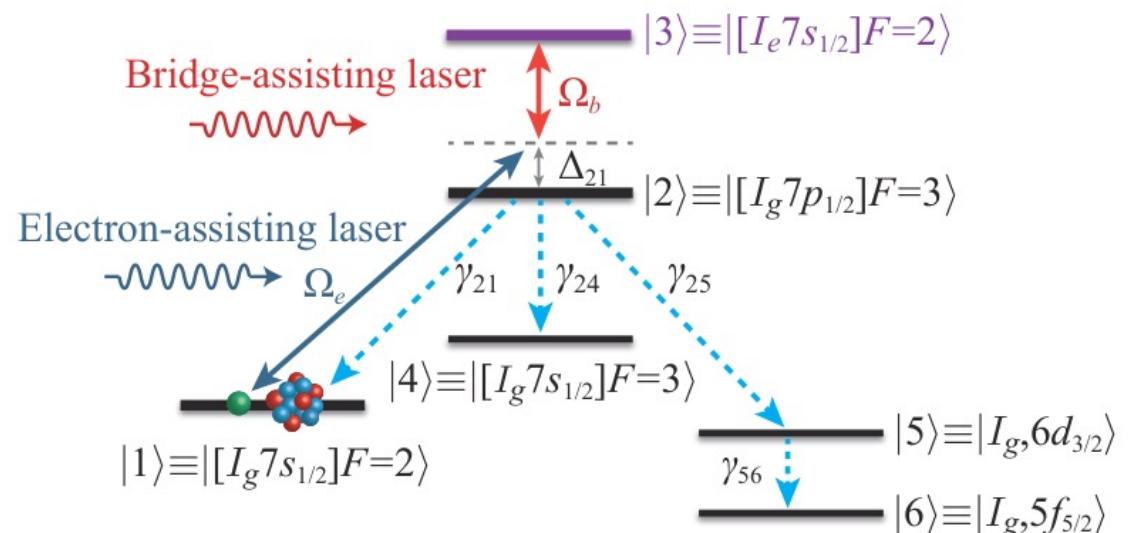
$$^{229}\text{Th}^{3+}: (I_g, 7p_{1/2}, F=3) \rightarrow (I_e, 7s_{1/2}, F=2)$$

$|[I_g 7p_{1/2}]F_2\rangle \equiv |2\rangle$  is unstable (1.1 ns) with three decay channels

Quantum master equation:  $\dot{\rho} = -i[H_I(t), \rho] + \mathcal{L}\rho$

$$H_I(t) = \Delta_{21}\hat{\sigma}_{22} + (\Omega_e(t)\hat{\sigma}_{21} + \Omega_b(t)\hat{\sigma}_{32} + \text{H.c.})$$

- $\Delta_{21}$  one-photon detuning
- $\Omega_e, \Omega_b$  Rabi frequencies



A schematic diagram of a six-level HEB scheme in  $^{229}\text{Th}^{3+}$

W. Wang, F. Zou, S. Fritzsche, and Y. Li, Phys. Rev. Lett. **133**, 223001 (2024)

## 2. Isomeric population transfer of $^{229}\text{Th}$

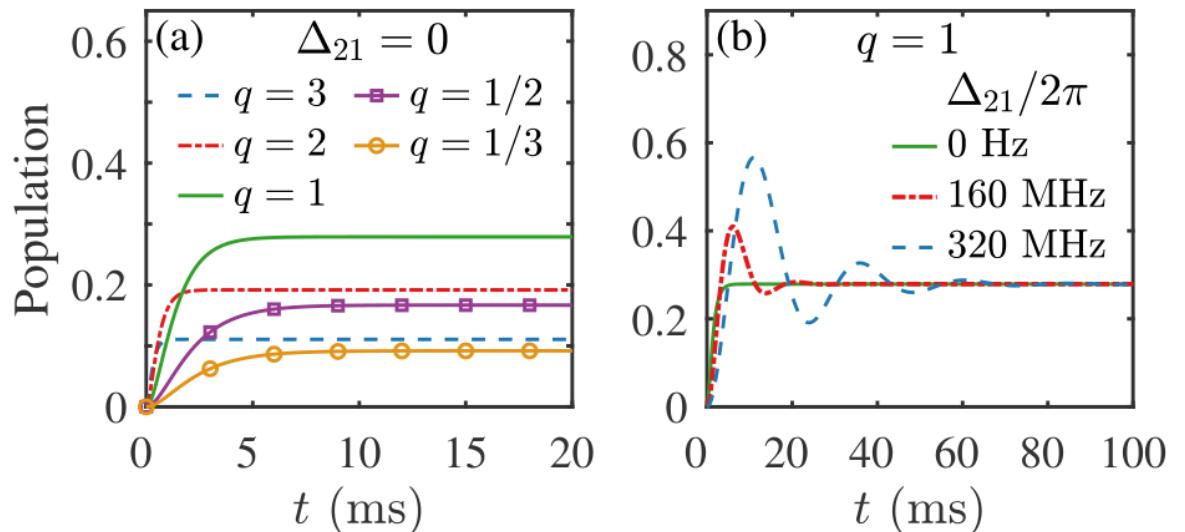
Time-independent scheme: envelop functions of Rabi frequencies  $f_b(t) = f_e(t) = 1$

$$q = |\Omega_e/\Omega_b|$$

Maximum isomeric population at  $q = 1$ : 27.9%

$\Delta_{21}$  does not affect the maximum isomeric population

Bridge-assisting laser intensity:  $3 \times 10^5 \text{ W/cm}^2$



(a, b) Isomeric population as a function of time

W. Wang, F. Zou, S. Fritzsche, and Y. Li, Phys. Rev. Lett. **133**, 223001 (2024)

## 2. Isomeric population transfer of $^{229}\text{Th}$

Time-independent scheme:

stimulated Raman adiabatic passage (STIRAP)

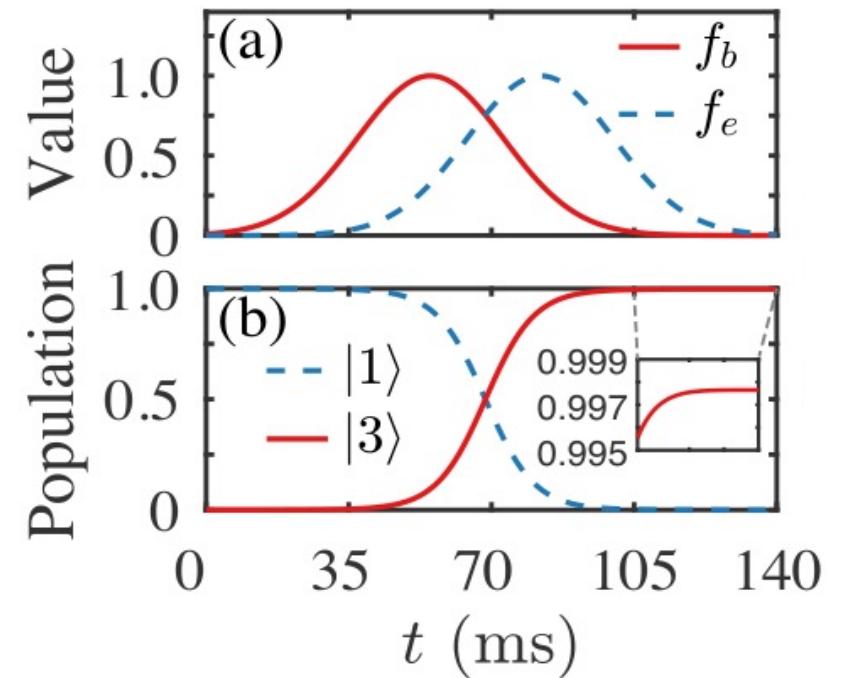
Envelop functions:  $f_b(t) = \exp[-(t-t_b)^2/T^2]$

$f_e(t) = \exp[-(t-t_e)^2/T^2]$

$|\Omega_b(t_b)| = |\Omega_e(t_e)|$  adiabatic condition :  $\Omega_b^2(t_b)T^2 \gg 1$

Bridge-assisting laser peak intensity:  $2 \times 10^7 \text{ W/cm}^2$

At final time 140 ms: isomeric population of 99.7%



(a) Gaussian envelop functions. (b) The population of  $|1\rangle$  and  $|3\rangle$  (isomeric state) as a function of time

**W. Wang, F. Zou, S. Fritzsche, and Y. Li, Phys. Rev. Lett. 133, 223001 (2024)**

### 3. The variation of fine-structure constant in HEB

Fine-structure constant:  $\alpha = \frac{e^2}{4\pi\epsilon_0\hbar c}$

Some theories beyond the Standard Model predict the possibility that the variation of  $\alpha$

Experiments: cosmological observations, atomic clock measurements

For atomic clock, introduce the sensitivity factor  $K_\alpha$ :  $\frac{\dot{\omega}}{\omega} = K_\alpha \frac{\dot{\alpha}}{\alpha}$ ,

For ordinary atomic clock,  $|K_\alpha| \sim 1$

$^{229}\text{Th}$  nuclear clock,  $K_\alpha = -8.2 \times 10^3$

P. Fadeev, J. C. Berengut, and V. V. Flambaum, Phys. Rev. A **102**, 052833 (2020)

### 3. The variation of fine-structure constant in HEB

In previous studies, electronic and nuclear transitions are treated as independent processes.

If correlation arises, how does this influence the clock performance?

HEB transition provides a natural framework to answer this question

$$K_\alpha = \frac{\Delta E_C}{\omega} \left[ 1 - \sum_t (|b_{i,t}|^2 + |b_{f,t}|^2) \right]$$

$\Delta E_C$  : the Coulomb energy difference between nuclear isomeric and ground states

$b_{i,t}$  and  $b_{f,t}$  are the mixing coefficients of the initial and final dressed hyperfine states. Since these mixing coefficients are small. Qualitatively, for a given nucleus, the smaller the transition energy, the larger the sensitivity factor.

**W. Wang, S. Fritzsche, and Y. Li, Phys. Rev. A 112, 022811 (2025)**

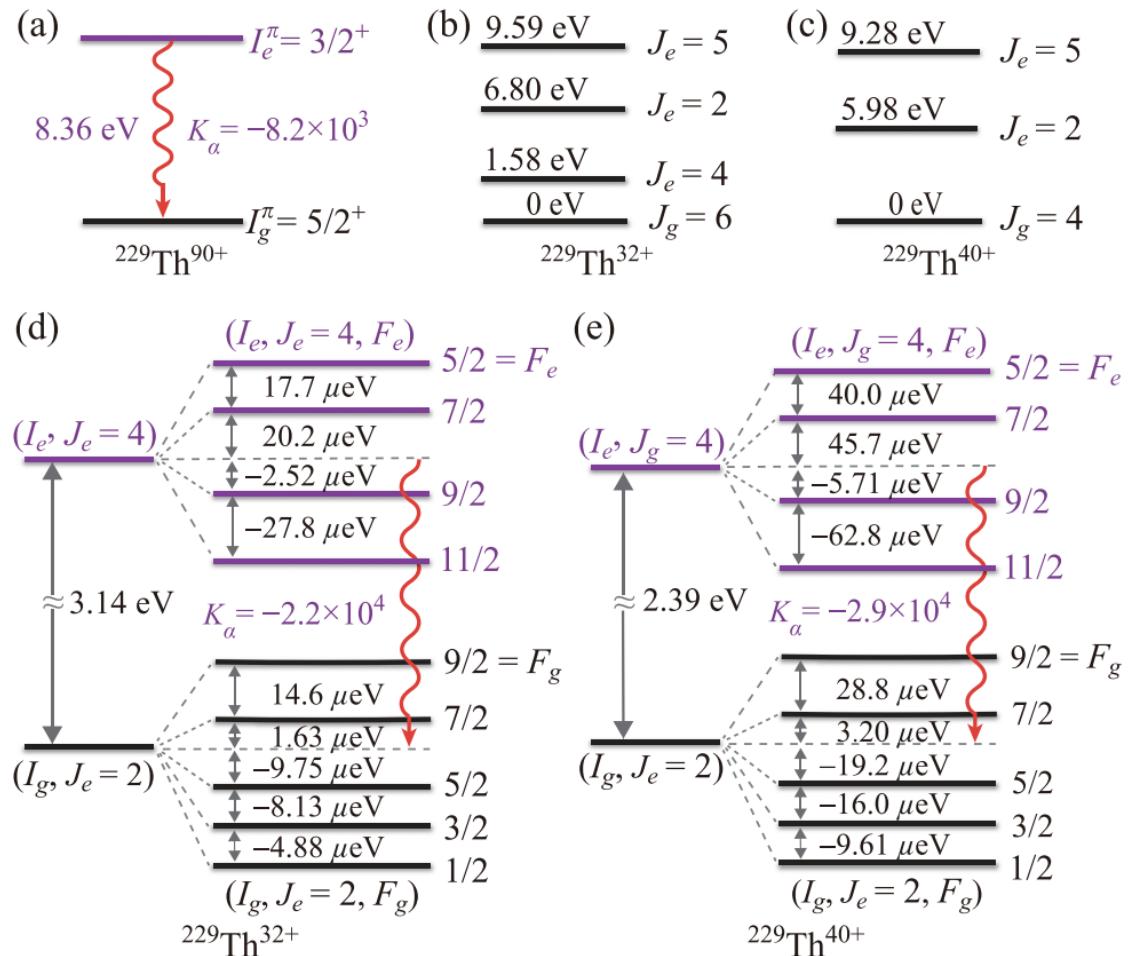
### 3. The variation of fine-structure constant in HEB

(a) Nuclear levels of  $^{229}\text{Th}^{90+}$ . (b) Partial electronic levels of  $^{229}\text{Th}^{32+}$ . (c) Partial electronic levels of  $^{229}\text{Th}^{40+}$ . (d) HEB transitions in  $^{229}\text{Th}^{32+}$ . (e) HEB transitions in  $^{229}\text{Th}^{40+}$

3.14-eV HEB transitions in  $^{229}\text{Th}^{32+}$

2.39-eV HEB transitions in  $^{229}\text{Th}^{40+}$

Compared with the bare nuclear transition,  $K_\alpha$  are enhanced by factors of 2.7 and 3.5, respectively, and the transitions lie in the visible spectrum.



W. Wang, S. Fritzsche, and Y. Li, Phys. Rev. A **112**, 022811 (2025)

### 3. The variation of fine-structure constant in HEB

- HEB transition rate formula

$$A_{\text{HEB}} = \frac{2[L][F_f]}{[L]!!^2} \left( \frac{\omega}{c} \right)^{[L]} \frac{L+1}{L} \times |\langle [I_f \gamma_f J_f] F_f | | \mathcal{O}^{(\tau L)} | | [I_i \gamma_i J_i] F_i \rangle|^2,$$

- A general formula for multipole transitions

$$\begin{aligned} & \langle [I_f \gamma_f J_f] F_f | | \mathcal{O}^{(\tau L)} | | [I_i \gamma_i J_i] F_i \rangle \\ & \approx \sum_t \left[ a_i b_{f,t}^* \begin{Bmatrix} J_t & J_i & L \\ F_i & F_f & I_i \end{Bmatrix} \langle \gamma_t J_t | | \mathcal{O}^{(\tau L)} | | \gamma_i J_i \rangle \right. \\ & \quad + a_f^* b_{i,t} \begin{Bmatrix} J_f & J_t & L \\ F_i & F_f & I_f \end{Bmatrix} \langle \gamma_f J_f | | \mathcal{O}^{(\tau L)} | | \gamma_t J_t \rangle \\ & \quad \left. \times (-1)^{J_t - J_i + I_f - I_i} \right]. \end{aligned}$$

**W. Wang, S. Fritzsche, and Y. Li, Phys. Rev. A **112**, 022811 (2025)**

# 3. The variation of fine-structure constant in HEB

Results:

TABLE I. The HEB transition rates for different channels in  $^{229}\text{Th}^{32+}$  ions.

Transition	Type	Rate (s <sup>-1</sup> )
$F_e = 5/2 \rightarrow F_g = 3/2$	$M1$	$1.3 \times 10^{-9}$
$F_e = 5/2 \rightarrow F_g = 5/2$	$M1$	$9.8 \times 10^{-9}$
$F_e = 5/2 \rightarrow F_g = 7/2$	$M1$	$4.4 \times 10^{-9}$
$F_e = 7/2 \rightarrow F_g = 5/2$	$M1$	$1.5 \times 10^{-9}$
$F_e = 7/2 \rightarrow F_g = 7/2$	$M1$	$3.7 \times 10^{-8}$
$F_e = 7/2 \rightarrow F_g = 9/2$	$M1$	$8.8 \times 10^{-9}$
$F_e = 9/2 \rightarrow F_g = 7/2$	$M1$	$1.2 \times 10^{-10}$
$F_e = 9/2 \rightarrow F_g = 9/2$	$M1$	$6.5 \times 10^{-8}$
$F_e = 11/2 \rightarrow F_g = 9/2$	$M1$	$2.0 \times 10^{-8}$

HEB transition rates  $\sim 10^{-9} \text{ s}^{-1}$

These HEB transitions can be controlled by the precision spectroscopy

TABLE II. Same as Table I, but for  $^{229}\text{Th}^{40+}$  ions.

Transition	Type	Rate (s <sup>-1</sup> )
$F_e = 5/2 \rightarrow F_g = 3/2$	$M1$	$9.1 \times 10^{-10}$
$F_e = 5/2 \rightarrow F_g = 5/2$	$M1$	$2.1 \times 10^{-9}$
$F_e = 5/2 \rightarrow F_g = 7/2$	$M1$	$5.7 \times 10^{-10}$
$F_e = 7/2 \rightarrow F_g = 5/2$	$M1$	$2.1 \times 10^{-9}$
$F_e = 7/2 \rightarrow F_g = 7/2$	$M1$	$5.5 \times 10^{-9}$
$F_e = 7/2 \rightarrow F_g = 9/2$	$M1$	$8.7 \times 10^{-10}$
$F_e = 9/2 \rightarrow F_g = 7/2$	$M1$	$7.6 \times 10^{-10}$
$F_e = 9/2 \rightarrow F_g = 9/2$	$M1$	$6.0 \times 10^{-9}$
$F_e = 11/2 \rightarrow F_g = 9/2$	$M1$	$4.0 \times 10^{-9}$

For the  $E3$  clock transition of the  $^{171}\text{Yb}^+$  atomic clock, its transition rate is  $\sim 10^{-9} \text{ s}^{-1}$

W. Wang, S. Fritzsche, and Y. Li, Phys. Rev. A **112**, 022811 (2025)

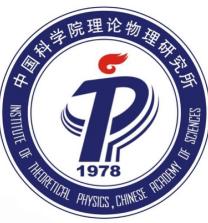
## 4. Summary and outlook

Summary:

1. The HEB transition is proposed and investigated. A quantum-optical model is constructed. By employing quantum control techniques, efficient isomeric population transfer is predicted.
2. In  $^{229}\text{Th}^{32+}$  and  $^{229}\text{Th}^{40+}$ , HEB transitions lie in the visible spectrum, and are more sensitive to the variation of the fine-structure constant than the bare nuclear transition.

Outlook:

1. By employing different quantum control techniques, the isomeric population transfer can be further optimized.
2. Since  $^{235}\text{U}$  has a nuclear excitation energy of about 76.7 eV, investigating HEB transitions in  $^{235}\text{U}$  is of interest.



Thank you for your attention!