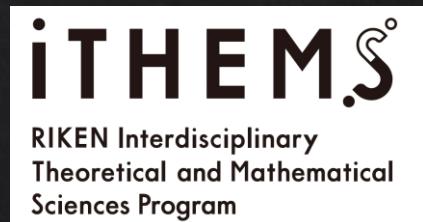


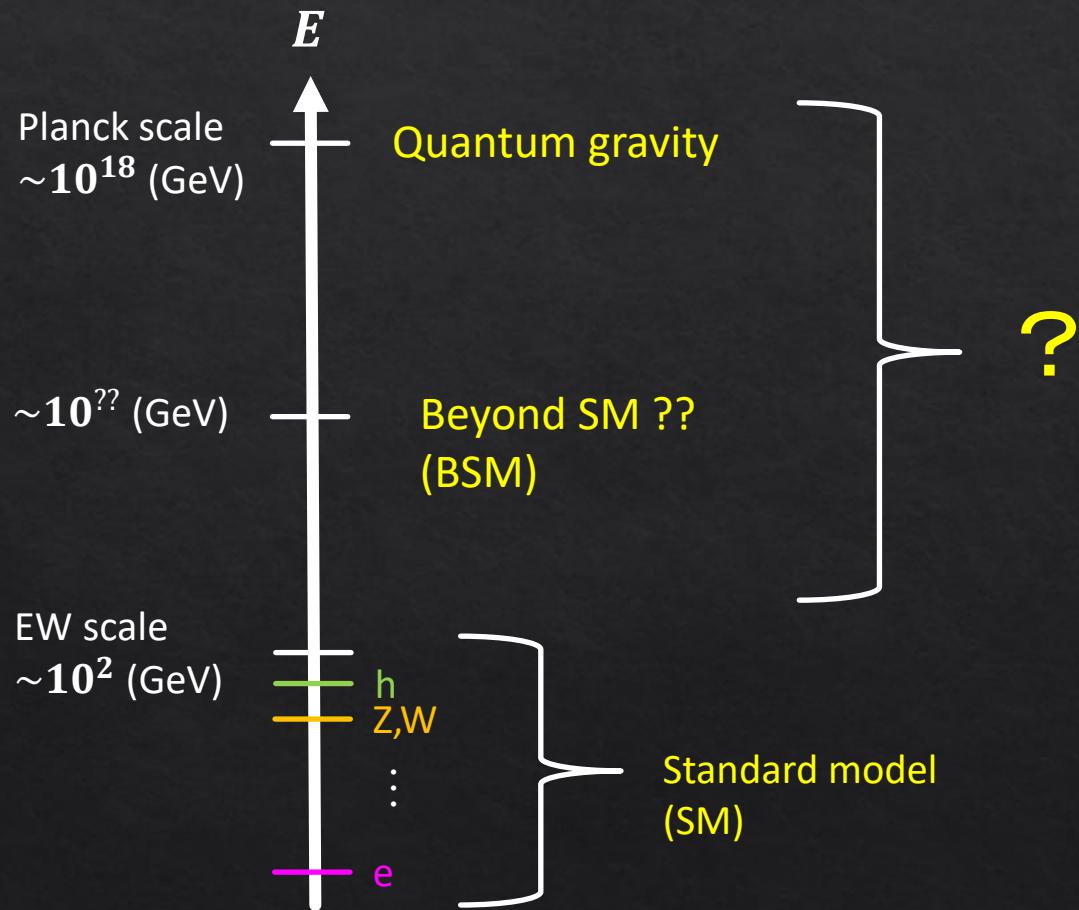
Probing New Physics through the Cosmological Collider

Shuntaro Aoki
(RIKEN iTHEMS)

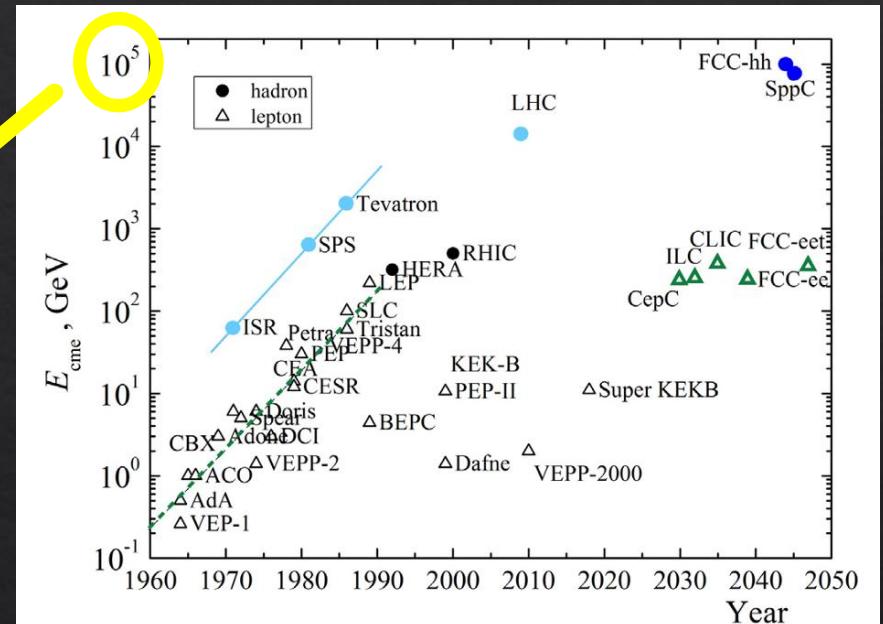
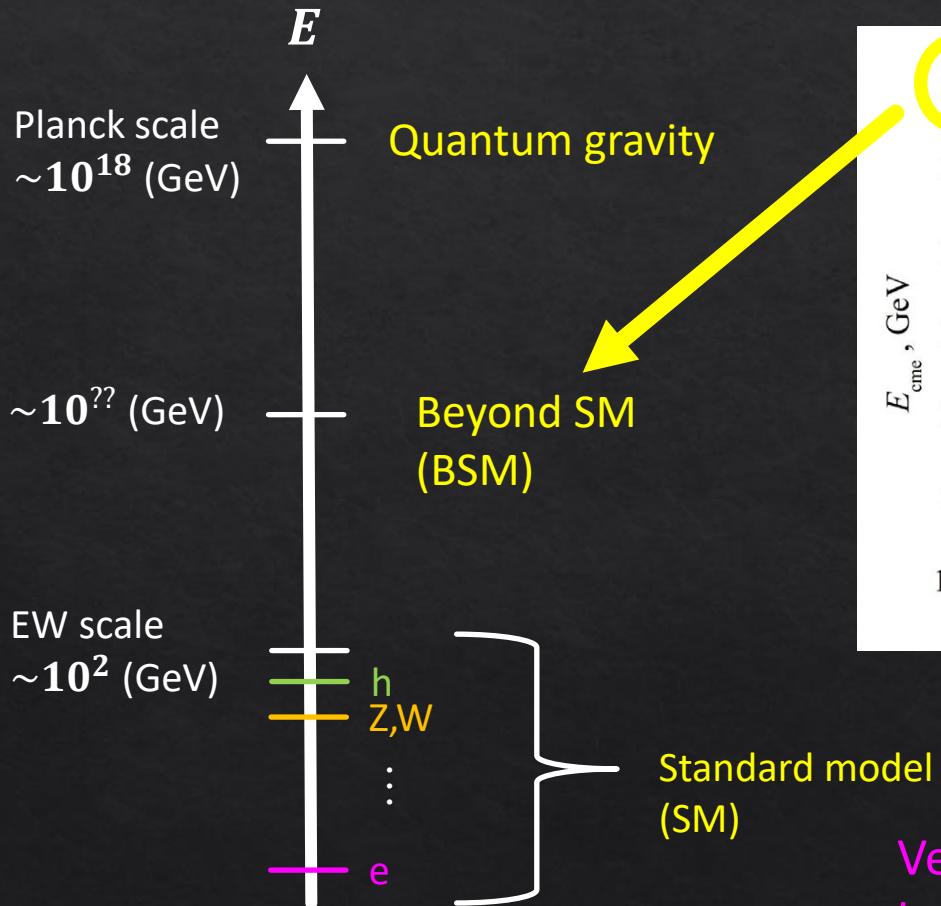


BPCS2025
September 26, 2025

Introduction



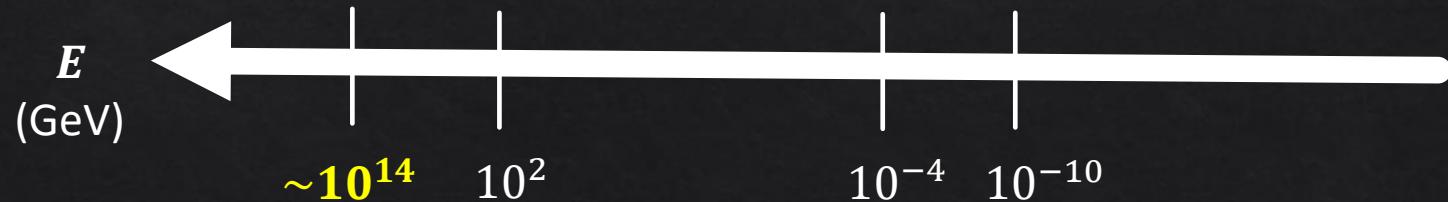
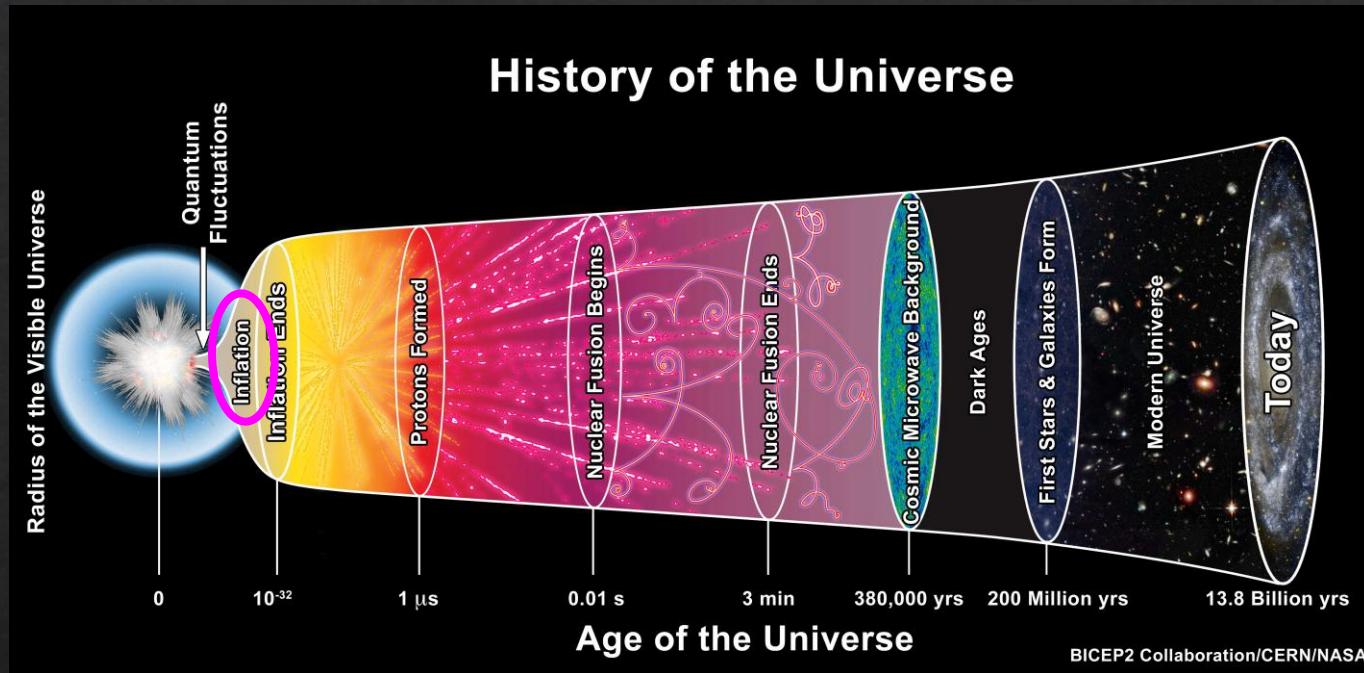
Introduction



Gray, Future colliders for the high-energy frontier

Very important,
but BSM might appear at higher scale
⇒ Another approach?

Universe as a gigantic accelerator

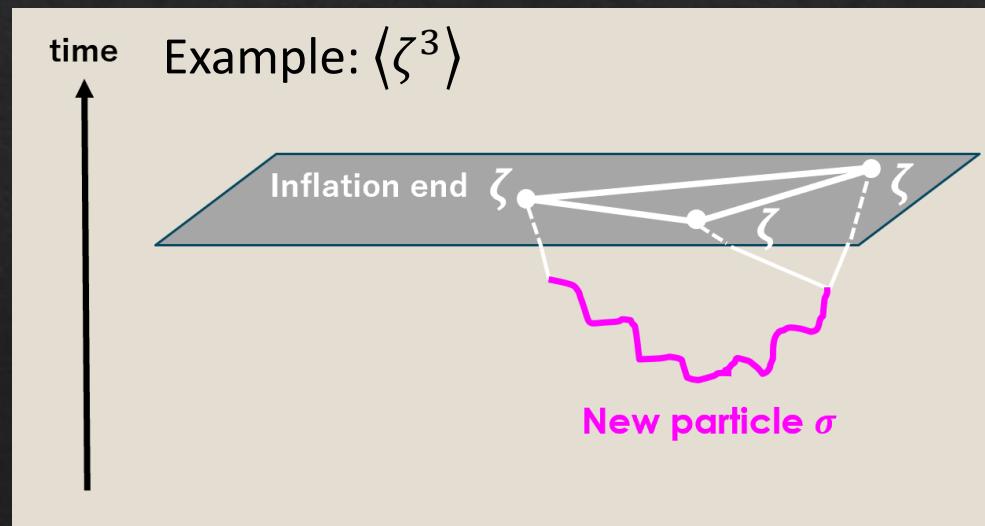


Idea: probing new physics with inflation energy

Cosmological Collider

Chen, Wang, '10
Baumann, Green, '12
Noumi, Yamaguchi, Yokoyama, '13
Arkani-Hamed, Maldacena, '15

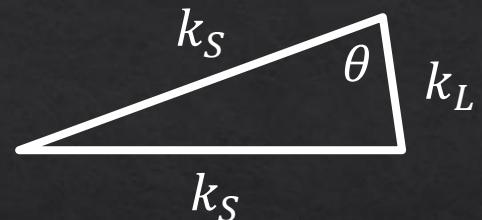
- $\langle \zeta^n \rangle$ can contain information of heavy (new) particle σ



Cosmic expansion creates σ with mass $\sim H$

Cosmological Collider

- Clear signal in squeezed limit of triangle $k_S \gg k_L$



- σ leaves specific imprint on $\langle \zeta^3 \rangle$

$$\langle \zeta^3 \rangle|_{k_L \ll k_S} \sim e^{-\pi\mu} \left(\frac{k_L}{k_s}\right)^{\frac{3}{2}} \cos \left[\mu \log \left(\frac{k_L}{k_s}\right) + \delta(\mu) \right] P_s(\cos \theta)$$

Boltzmann suppression

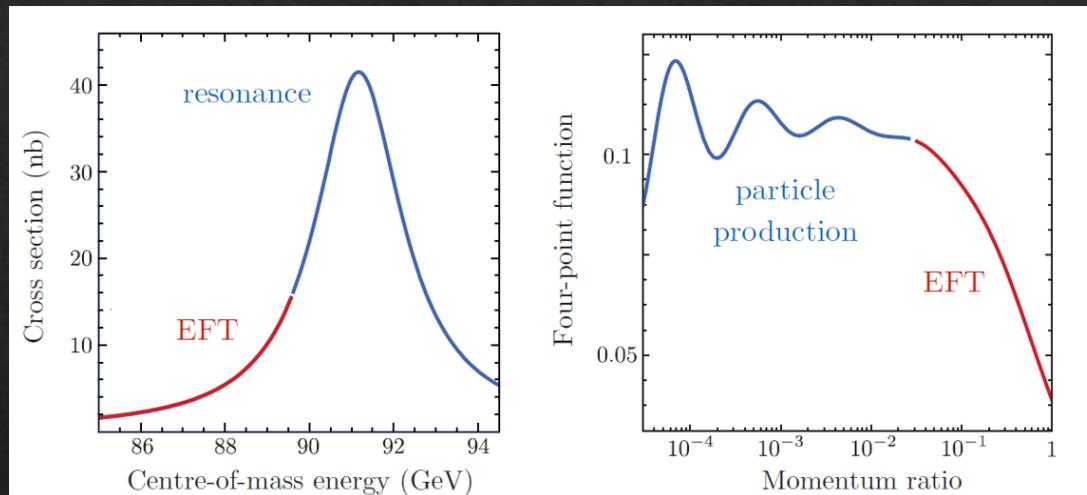
$\mu \sim m_\sigma/H$: mass

spin

dilution

Cosmological Collider

- Particle physics analogy

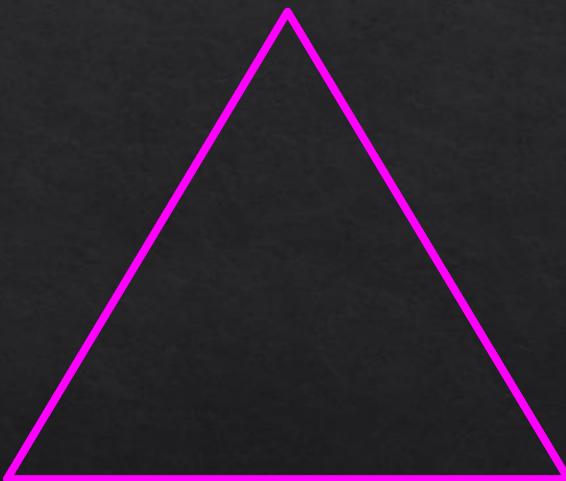


Baumann, Lectures on The Cosmological Bootstrap

- $m_\sigma \sim H_{\text{inf}} \sim 10^{14}(\text{GeV}) \ggg 10^{4,5}(\text{GeV}) \sim \text{collider on earth}$

Several directions

Understand
structure of correlators
(→ Xianyu's talk)



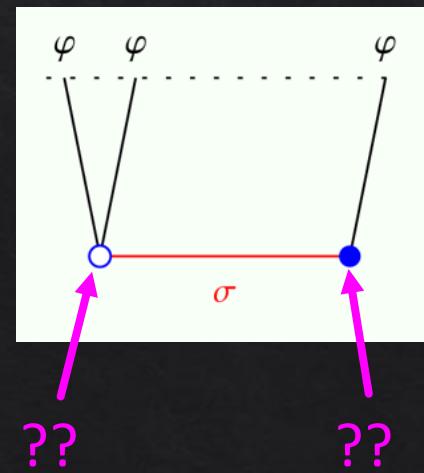
Comparison with
Observation
(→ Zhong's talk)

Application to BSM
(model building aspects)
(→ My talk)

Condition for large CC signal??

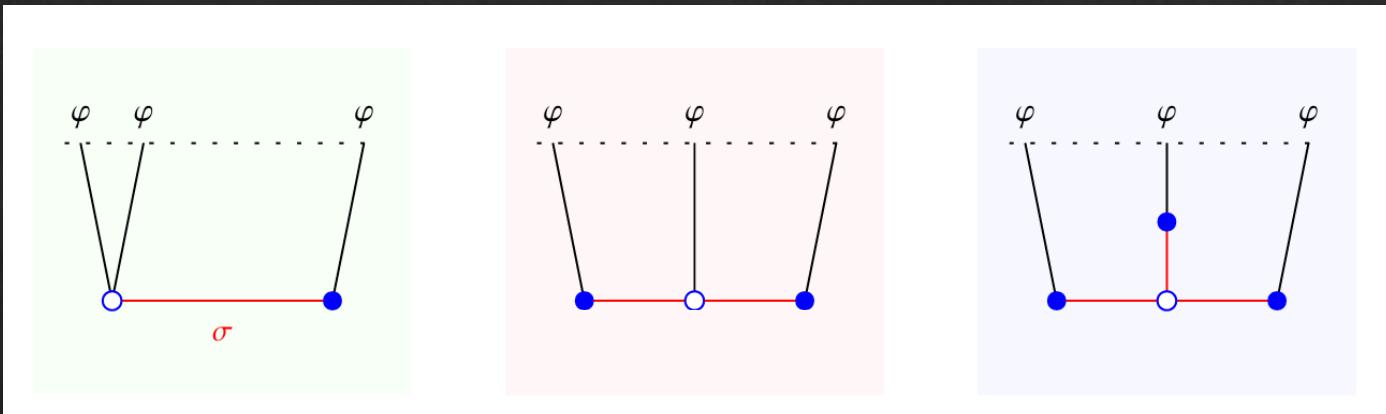
$$\langle \zeta^3 \rangle|_{k_L \ll k_s} \sim e^{-\pi\mu} \left(\frac{k_L}{k_s} \right)^{\frac{3}{2}} \cos \left[\textcolor{blue}{\mu} \log \left(\frac{k_L}{k_s} \right) + \delta(\mu) \right] P_s(\cos \theta)$$

??



Summary:

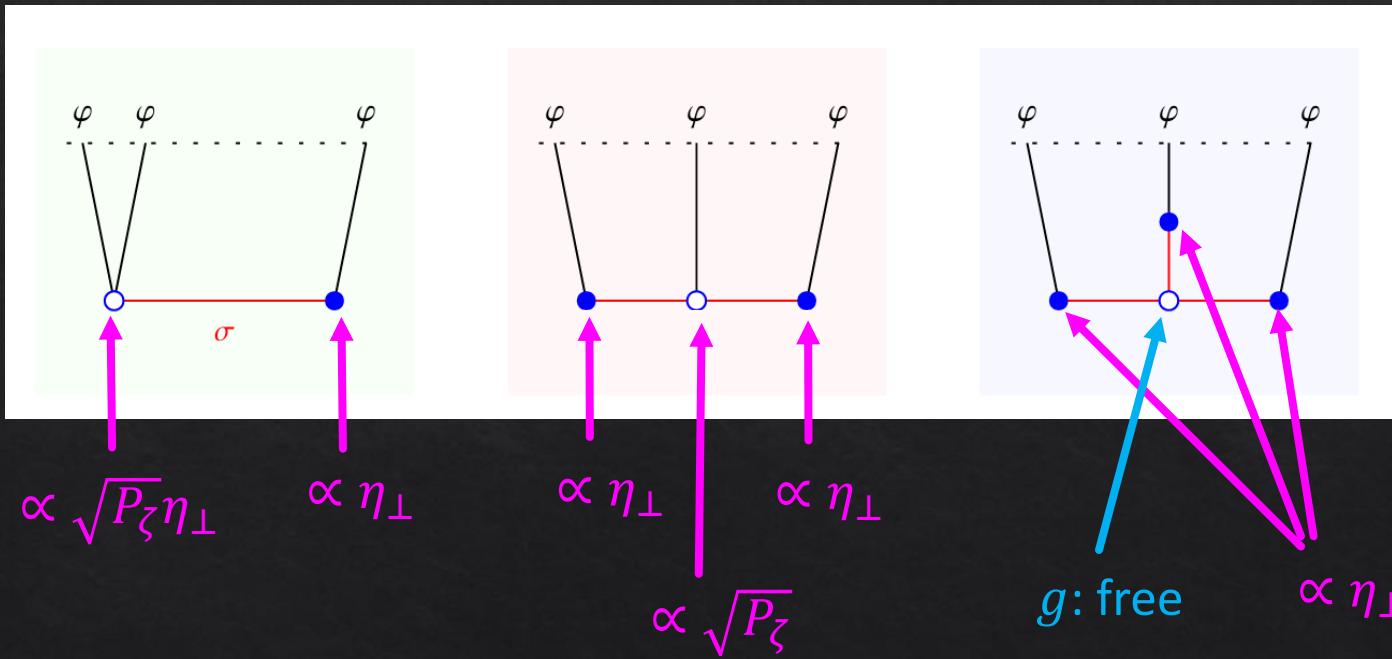
3 dominant diagram for $\langle \zeta^3 \rangle$



Summary: 3 dominant diagram for $\langle \zeta^3 \rangle$

Assume $m_\sigma \sim H$ (w/o fine-tuning)

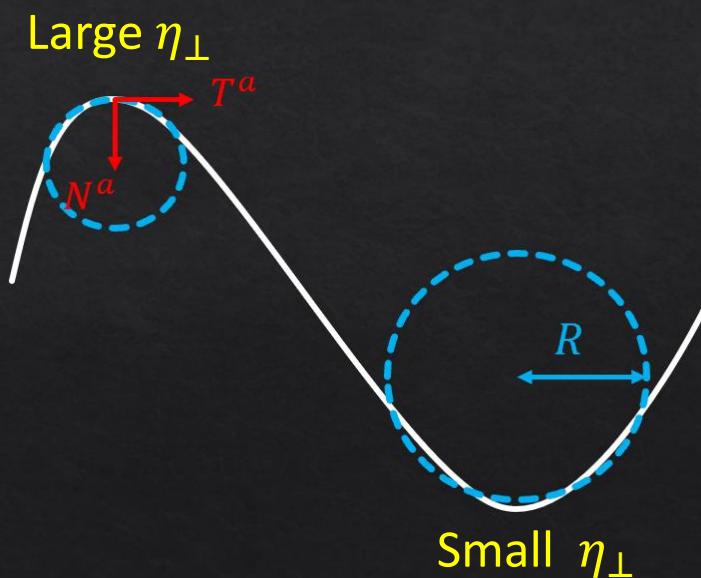
$\sqrt{P_\zeta} \sim 10^{-5}$: const



“Quasi single field inflation”
Chen, Wang, ’10

Turn

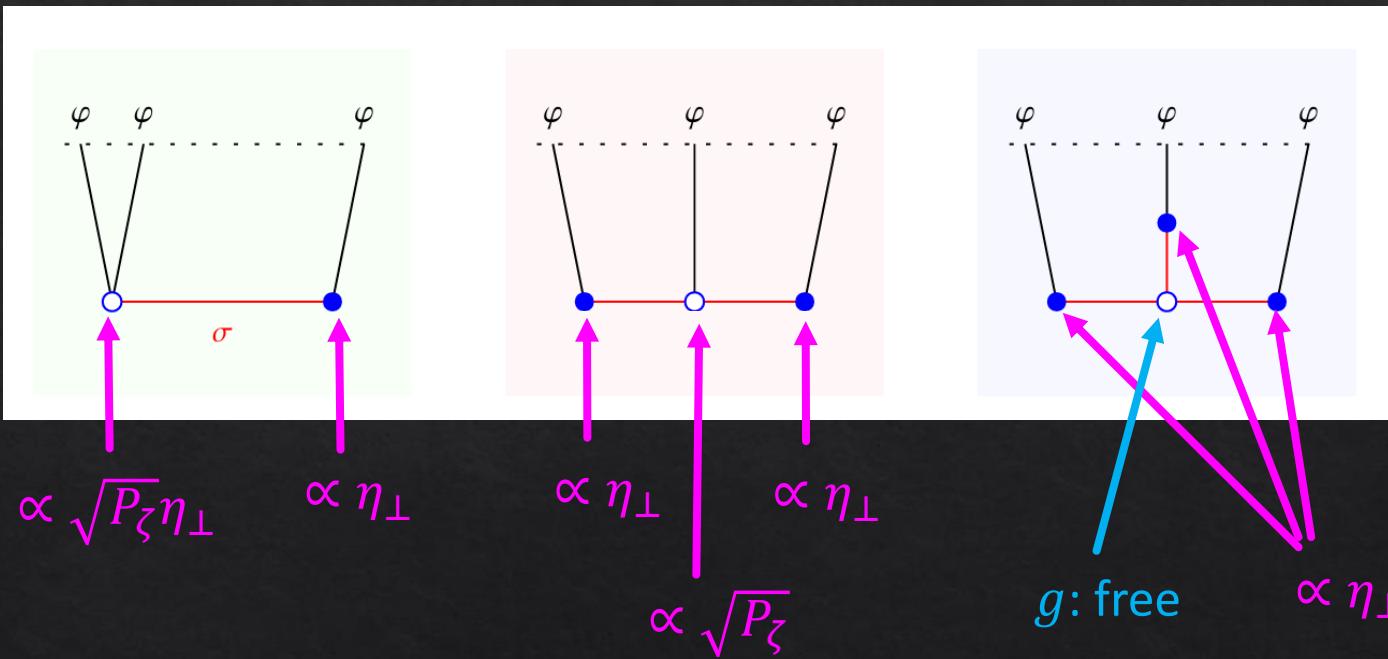
- Turn rate: $\eta_{\perp} = \frac{d\phi}{dN} \times \kappa$ (similar to $\omega = v/R$ in Newtonian mechanics)



Summary: 3 dominant diagram for $\langle \zeta^3 \rangle$

Assume $m_\sigma \sim H$ (w/o fine-tuning)

$\sqrt{P_\zeta} \sim 10^{-5}$: const



CC signal

$\propto \eta_\perp^2 (\times \text{oscillation} \times \text{BF})$

$\propto \eta_\perp^2$

$\propto \eta_\perp^3 \times g / H \sqrt{P_\zeta}$

Conditions for large signal

- Large turn rate η_{\perp} . Remember $\eta_{\perp} = \frac{d\phi}{d\mathcal{N}} \times \kappa = \sqrt{2\epsilon} M_{\text{PL}} \times \kappa$
⇒ Large $\kappa \Rightarrow$ sub-Planckian scale
- Large cubic σ^3 (triple exchange)

Check by concrete model

- ϕ (Higgs or dilaton) + R^2 (Scalaron) Small CC signal $\sim O(\epsilon)$

SA, A. Ghoshal and A. Strumia, 2408.07069

- Multifield α -attractor

SA, Roest, Werth, in progress

- Monodromy

SA, Otsuka, Yanagita, 2509.06739



Large $\sim O(1)$

Scalar $\phi + R^2$

SA, A. Ghoshal and A. Strumia,
2408.07069

➤ J-frame


$$S = \int d^4x \sqrt{|\det g|} \left[-\frac{1}{2}f(\phi)R + \frac{R^2}{6f_0^2} + \sum_{\phi} \frac{(D_{\mu}\phi)(D^{\mu}\phi)}{2} - V_J(\phi) \right]$$

- ϕ : some scalars (specified later)
- R^2 gives an additional scalar z (scalaron)

➤ E-frame

$$S = \int d^4x \sqrt{|\det g|} \left[-\frac{\bar{M}_{\text{Pl}}^2}{2}R + \frac{6\bar{M}_{\text{Pl}}^2}{z^2} \frac{(\partial_{\mu}z)^2 + \sum_{\phi}(D_{\mu}\phi)^2}{2} - V(\phi, z) \right]$$

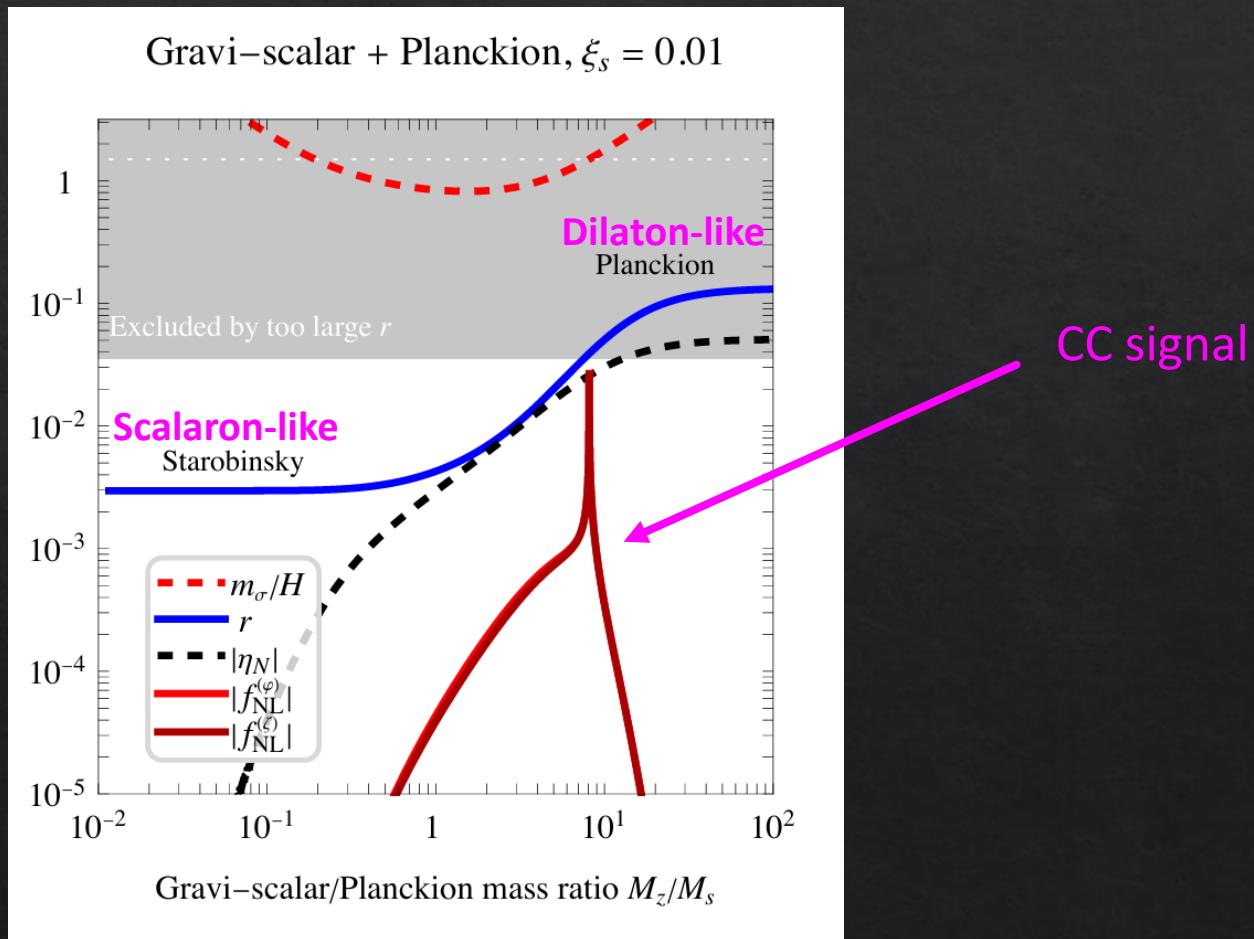
$$V = \left(\frac{6\bar{M}_{\text{Pl}}^2}{z^2} \right)^2 \left[V_J(\phi) + \frac{3}{8}f_0^2 (f + \xi_z z^2)^2 \right]. \quad \xi_z = -1/6,$$

Include two important models

- $\phi = \text{Dilaton} : f = \xi\phi^2, V_J = \lambda(\phi)\phi^4/4$ Kannike et al, '15, ...
 ← Planck scale is induced by $\langle\phi^2\rangle = M_{\text{pl}}^2/\xi_s$
- $\phi = \text{Higgs} : f = M_{\text{PL}}^2 + \xi\phi^2, V_J = \lambda\phi^4/4$ Salvio, Mazumdar '15, Ema '17, ...
 ← Unitarizing Higgs inflation $\xi f_0^2 \sim O(1)$

Only scale $\sim M_{\text{PL}}$ for natural parameter choice
 $\Rightarrow \kappa \propto 1/M_{\text{Pl}} \Rightarrow \eta_N \propto \sqrt{\epsilon}$: Small turn !!

CC signal from dilaton+ R^2



Multifield α -attractor

- Two field generalization of α -attractor ($\alpha, R_m \equiv m_\chi^2/m_\phi^2$)

$$G_{IJ} = \frac{6\alpha}{(1 - \phi^2 - \chi^2)^2} \delta_{IJ} .$$

$$V(\phi, \chi) = \frac{\alpha}{2} (m_\phi^2 \phi^2 + m_\chi^2 \chi^2) ,$$

- $\phi = r \cos \theta, \chi = r \sin \theta$

- radial (r) inflation for $\alpha \gg 1$ (~single α -attractor)

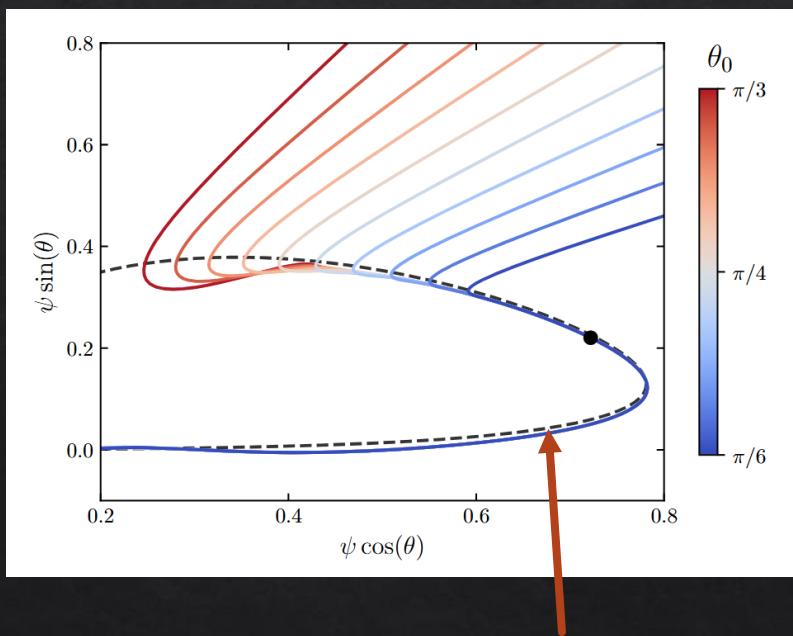
Kallosh, Linde, Roest, 1311.0472

- Angular (θ) inflation for $\alpha \ll 1$

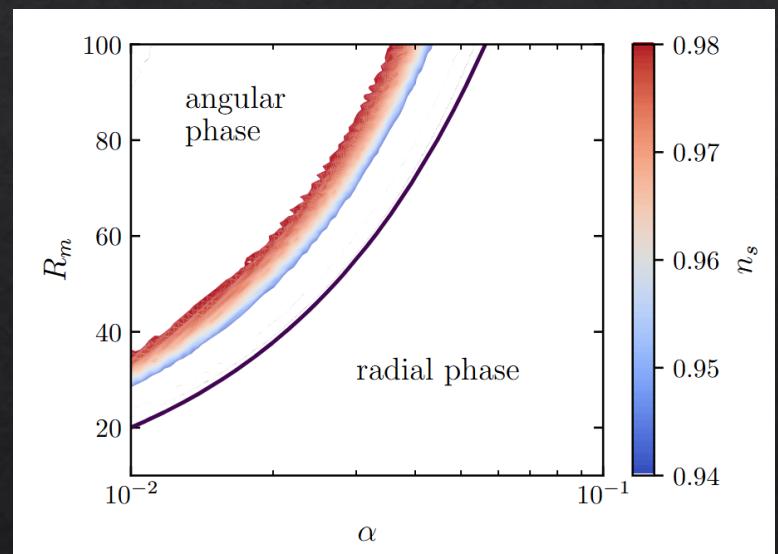
Christodoulidis, Roest, Sfakianakis, 1803.09841

Multifield α -attractor

Trajectory (attractor)



Parameter space



$$1 - r^2(\theta) \approx \frac{9\alpha(\cot \theta + R_m \tan \theta)^2}{2(R_m - 1)^2}$$

Multifield α -attractor

- Large turn rate: $\eta_{\perp} \sim 1$ for $\alpha \sim \epsilon \ll 1$ (curvature of trajectory $\kappa \propto 1/\alpha M_{\text{Pl}}$)

SA, Roest, Werth, in progress

$$\eta_{\perp}^2 \simeq \frac{4\epsilon}{3\alpha}$$

- isocurvature mass

$$m_{\sigma, \text{eff}}^2 = \eta_{\perp}^2 H^2$$

$m_{\sigma, \text{eff}} \sim H$ for $\eta_{\perp} \sim 1$
 $\Rightarrow O(1)$ CC signal

Monodromy

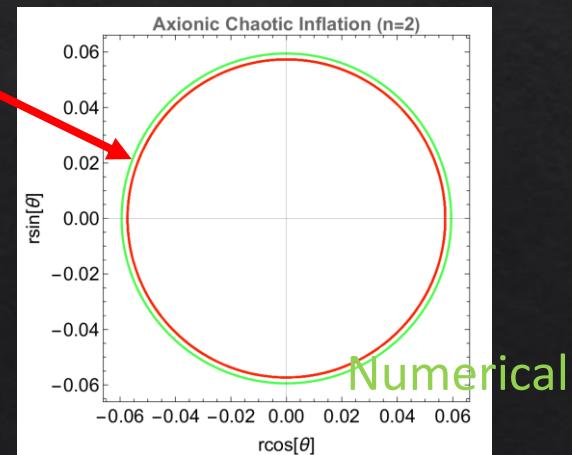
SA, Otsuka, Yanagita, 2509.06739

$$\left[\begin{array}{l} \text{Flat target space } G_{IJ} = \text{diag}(1, r^2) \\ \\ V(r, \theta) = \alpha \theta^n + \frac{M^2}{2} (r - R)^2, \end{array} \right]$$

Trajectory with $\frac{R}{r} \ll 1$:
$$\frac{n^2 \alpha^2 \theta^{2n-2}}{r^4} \simeq 3m^2 (m^2 r^2 + \alpha \theta^n)$$

Turn rate:
$$\eta_\perp^2 \simeq \frac{2\epsilon M_{\text{Pl}}^2}{r^2}$$
 curvature $\kappa \propto 1/r$
 \Rightarrow large $\eta_\perp \Rightarrow$ large CC

iso mass:
$$m^2 = \frac{M^2 R}{r}.$$

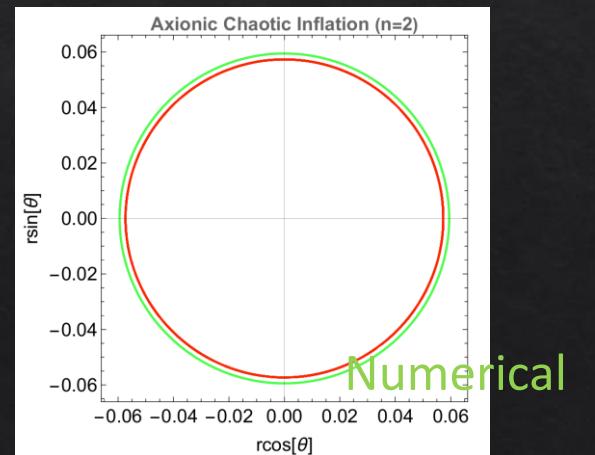


Monodromy

SA, Otsuka, Yanagita, 2509.06739

$$\left[\begin{array}{l} \text{Flat target space } G_{IJ} = \text{diag}(1, r^2) \\ V(r, \theta) = \alpha \theta^n + \frac{M^2}{2} (r - R)^2, \end{array} \right]$$

Parameters: (n, α, M, R)	n_s	r	$ \eta_\perp \equiv \rho /(2H)$	m/H
$\left(\frac{1}{2}, 2.7 \times 10^{-11}, 3 \times 10^{-4}, 4 \times 10^{-4}\right)$	0.98	0.03	16.2	4.77
$(1, 6.1 \times 10^{-13}, 1.5 \times 10^{-4}, 4 \times 10^{-4})$	0.975	0.018	10.5	2.15
$(2, 1 \times 10^{-14}, 3.5 \times 10^{-5}, 8 \times 10^{-4})$	0.967	0.022	2.16	0.85

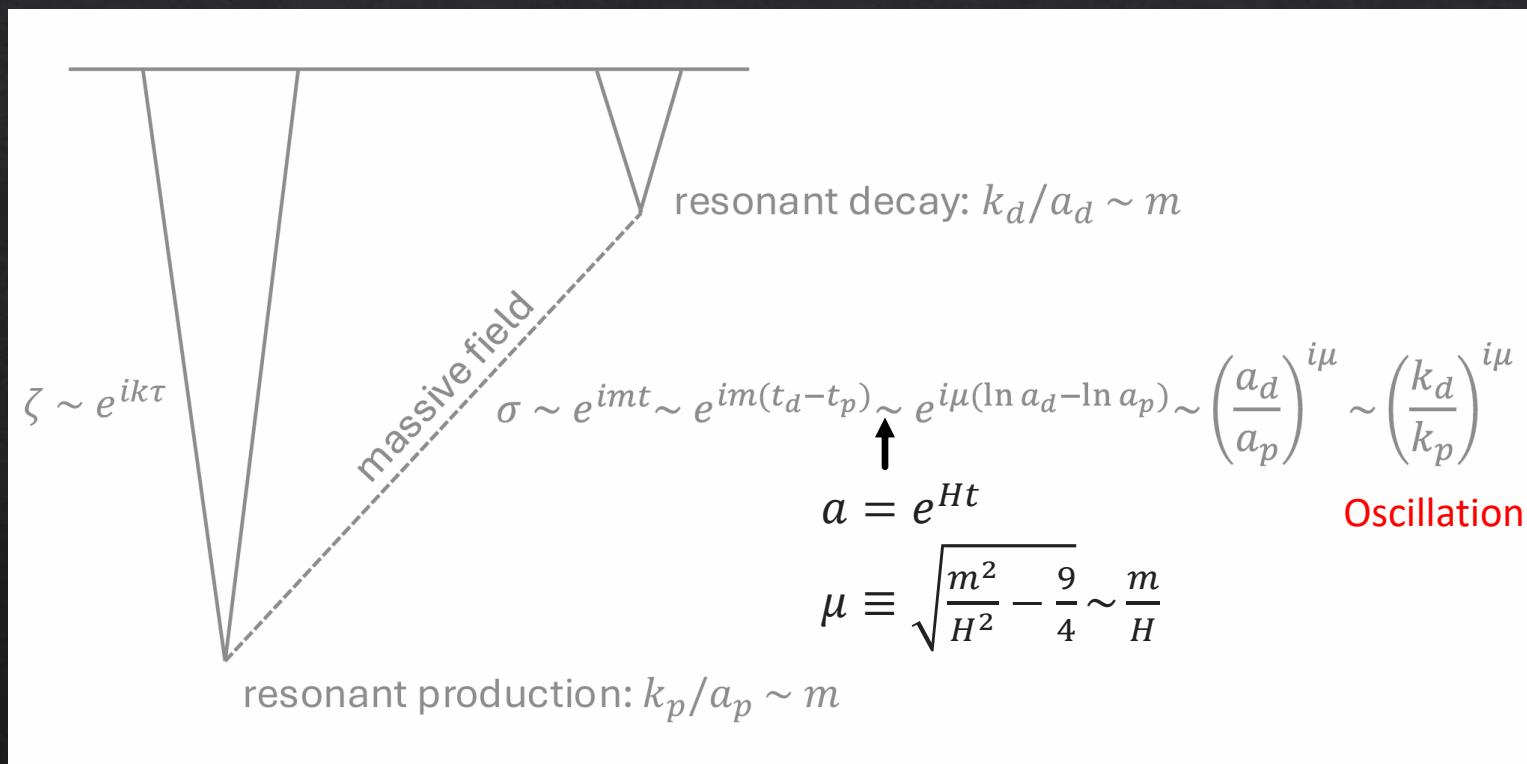


Summary

- Model building aspects of CC with tree scalar exchange (conditions for large signal, explicit models)
- Need big turn rate and/or big σ^3 for large signal
 - ↓
$$\eta_{\perp} = \sqrt{2\epsilon} M_{\text{PL}} \times \kappa \Rightarrow \text{Large } \kappa$$
- Three models
 - $\phi + R^2 \Rightarrow \kappa \propto 1/M_{\text{Pl}}$
 - Multifield α -attractor $\Rightarrow \kappa \propto 1/\alpha M_{\text{Pl}}$
 - Monodromy $\Rightarrow \kappa \propto 1/r$

Thank you!!

Why oscillation??



From Yi Wang's slide

Target & Assumption

- $\langle \zeta^3 \rangle$
- Inflation scenario
- Tree \gg Loop
- Scalar exchange without chemical potential

Many interesting possibilities if relax assumptions

Modulated reheating (Lu, Wang, Xianyu, 1907.07390)

Curvaton (Kumar, Sundrum, 1908.11378)

Chemical potential (Bodas, Kumar, Sundrum, 2010.04727)

Non-BD vacuum (Yin, 2309.05244)

...

From J to E-frame

➤ J-frame

$$S = \int d^4x \sqrt{|\det g|} \left[-\frac{1}{2}f(\phi)R + \frac{R^2}{6f_0^2} + \sum_{\phi} \frac{(D_{\mu}\phi)(D^{\mu}\phi)}{2} - V_J(\phi) \right]$$

- Add 0 (χ : auxiliary field) $-(R + 3f_0^2\chi/2)^2/6f_0^2$ which cancels R^2
- Do Weyl rescaling $g_{\mu\nu}^E = g_{\mu\nu}(f + \chi)/\bar{M}_{\text{Pl}}^2$
- χ becomes dynamical, define $z \equiv \sqrt{6(f + \chi)}$

➤ E-frame

$$S = \int d^4x \sqrt{|\det g|} \left[-\frac{\bar{M}_{\text{Pl}}^2}{2}R + \frac{6\bar{M}_{\text{Pl}}^2}{z^2} \frac{(\partial_{\mu}z)^2 + \sum_{\phi}(D_{\mu}\phi)^2}{2} - V(\phi, z) \right]$$

$$V = \left(\frac{6\bar{M}_{\text{Pl}}^2}{z^2} \right)^2 \left[V_J(\phi) + \frac{3}{8}f_0^2 (f + \xi_z z^2)^2 \right]. \quad \xi_z = -1/6,$$