



東南大學
SOUTHEAST UNIVERSITY



Classical Gravitational Dynamics from Quantum Field Theory

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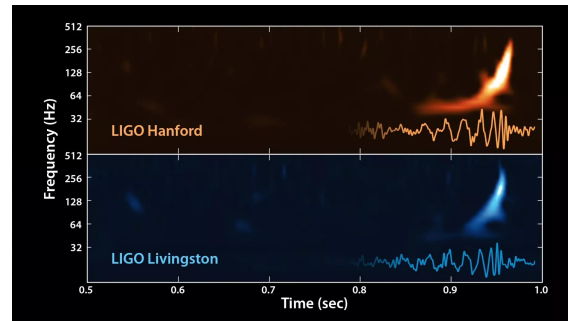
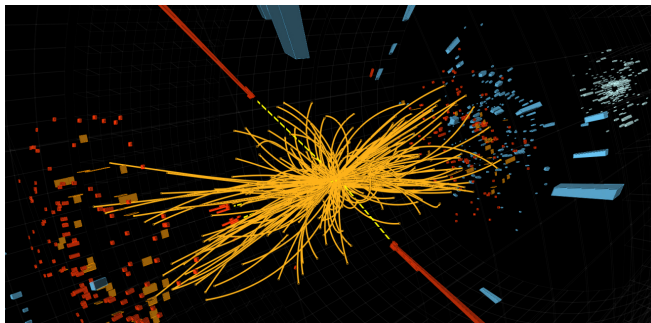
第十四届新物理研讨会

济南 2025.07.24



Precision era of fundamental physics

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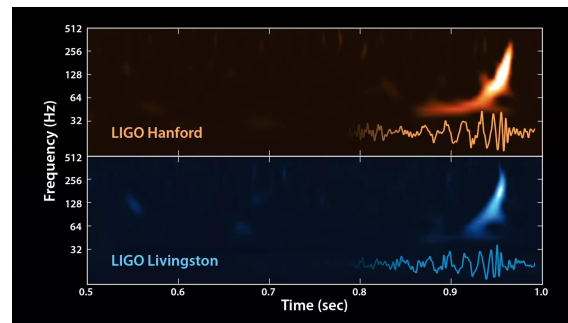
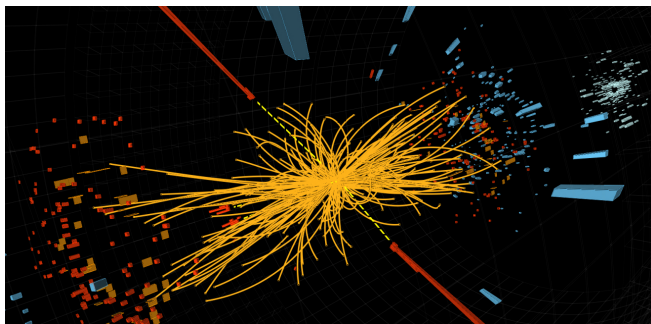
Two historic breakthroughs in science:

- Higgs bosons at the LHC (2012)
- Gravitational waves by the LIGO (2016)
- High-energy and gravitational physics entered a precision era!

Opened a new window on the Universe!

Precision era of fundamental physics

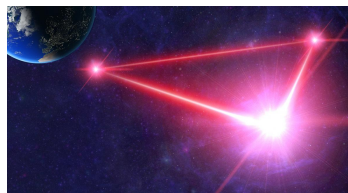
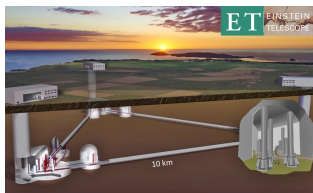
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Two historic breakthroughs in science:

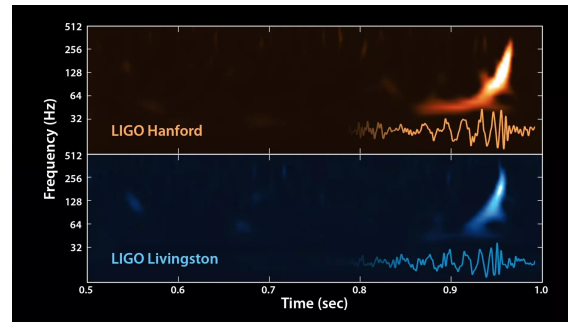
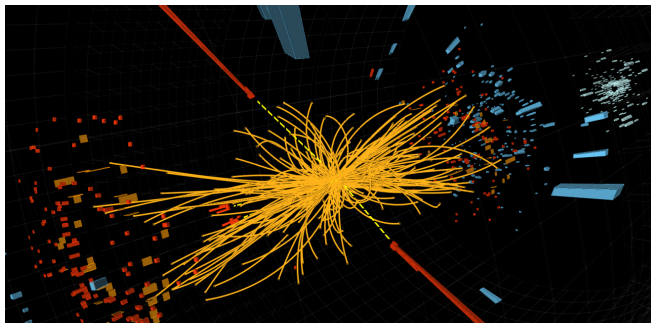
- Higgs bosons at the LHC (2012)
- Gravitational waves by the LIGO (2016)
- High-energy and gravitational physics entered a precision era!

Opened a new window on the Universe! **discovery potential** = **precise theoretical predictions!**



Precision era of fundamental physics

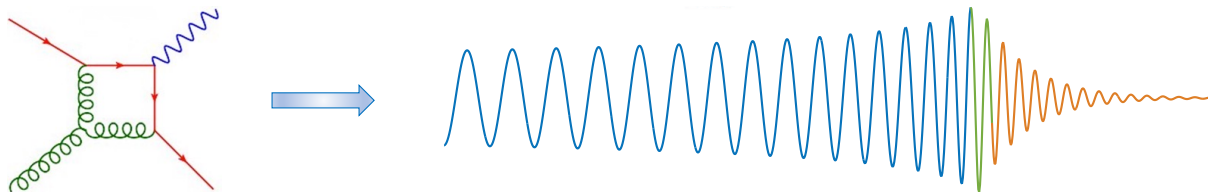
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Two historic breakthroughs in science:

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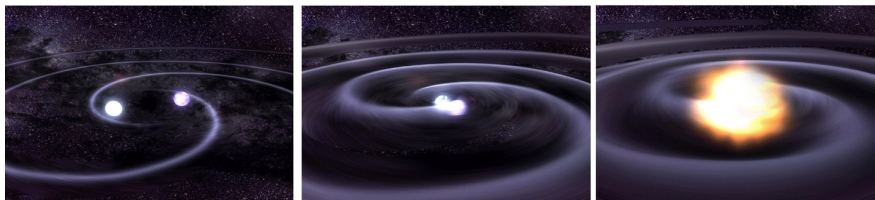
Modern techniques from LHC physics are playing a crucial role in precision GW physics!



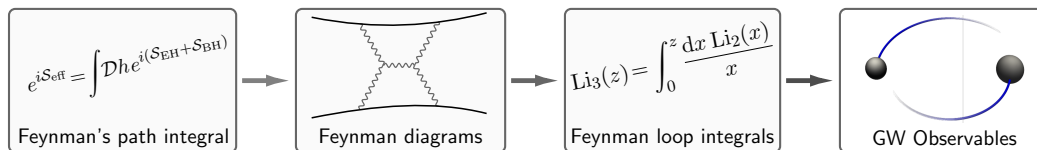
Outline

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- Introduction: gravitational two-body problems



- An Effective Field Theory for two-body dynamics



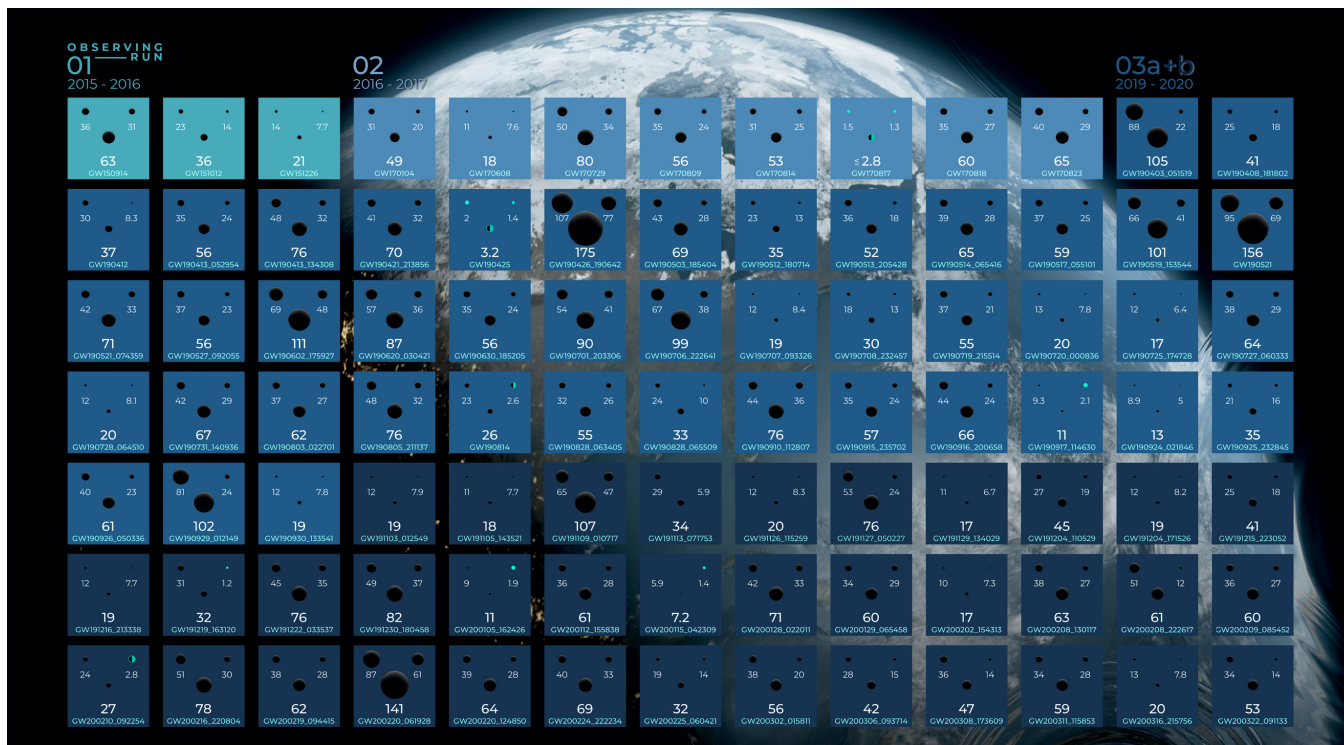
- Gravitational binary dynamics to NNNLO

$$\alpha_1 G + \alpha_2 G^2 + \alpha_3 G^3 + \alpha_4 G^4 + \alpha_5 G^5 + \dots$$

- Conclusion & Outlook

Gravitational waves from binary coalescences

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Credit: Carl Knox (OzGrav, Swinburne)

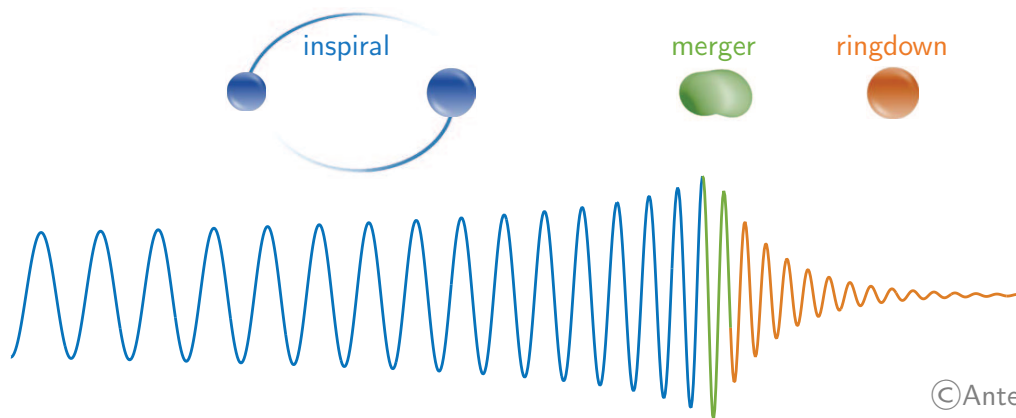


GWTC-3: 90 GW events—the majority are binary black holes (BH), but also several binary neutron stars (NS) and mixed NS-BHs.

gwosc.org

Gravitational waves from binary coalescences

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©Antelis & Moreno 2016

Merger: Numerical Relativity

Ringdown: black hole perturbation theory

Inspiral: the interaction between two bodies is weak

$$v^2 \sim \frac{GM}{r} \ll 1$$

- Numerical Relativity: accurately, but computationally expensive
- Analytic methods: corrections in v or G are perturbatively calculable

Post-Newtonian/post-Minkowskian expansion

► LHC theory technology, QFT methodology, shown great power!

Effective Field Theory

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- Gravitational binary system

$$S_{\text{WL}} = \sum_{i=1,2} \left[-\frac{m_i}{2} \int dt g_{\mu\nu} \dot{x}_i^\mu \dot{x}_i^\nu + \dots \right] \quad S_{\text{GR}} = \frac{-1}{16\pi G} \int d^4x \sqrt{-g} R + \dots$$

- Effective action for gravitational binary systems

Goldberger-Rothstein hep-th/0409156

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$e^{iS_{\text{eff}}[x_a(\tau)]} = \int \mathcal{D}h_{\mu\nu} e^{iS_{\text{WL}} + iS_{\text{GR}}}$$

- Post-Minkowskian expand in powers of G

$$L_{\text{eff}} = L_0 + GL_1 + G^2L_2 + \dots \quad L_0 = -\sum_i \frac{m_i}{2} \eta_{\mu\nu} \dot{x}_i^\mu \dot{x}_i^\nu$$

- The equations of motion for trajectories:

Kälín-Porto 2006.01184

$$m_i \ddot{x}_i^\mu = -\eta^{\mu\nu} \sum_{n=1}^{\infty} G^n \left(\frac{\partial L_n}{\partial x_i^\nu} - \frac{d}{d\tau_i} \frac{\partial L_n}{\partial \dot{x}_i^\nu} \right) \quad x_i^\mu = b_i^\mu + u_i^\mu \tau_i + \delta x_i^\mu(\tau_i) + \dots$$

- Physical observables:

$$\Delta p_i^\mu = p_i^\mu(+\infty) - p_i^\mu(-\infty) = -\eta^{\mu\nu} \sum_{n=1}^{\infty} G^n \int_{-\infty}^{\infty} d\tau_i \left(\frac{\partial L_n}{\partial x_i^\nu} \right)$$

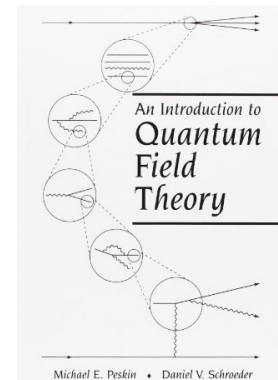
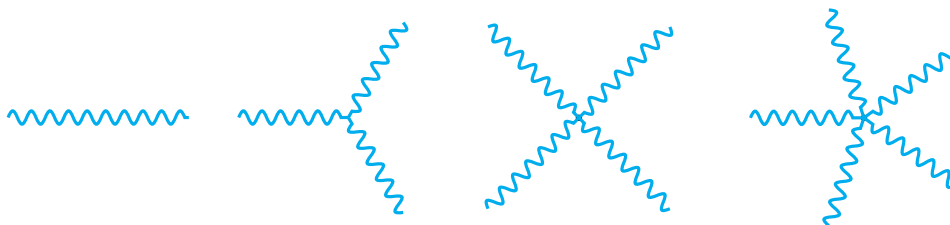
Effective Field Theory

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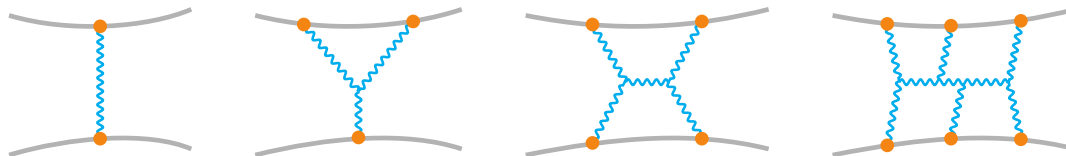
- Worldlines as classical sources in path integral:



- Hilbert-Einstein: $\mathcal{L}_{\text{HE}} = \mathcal{L}_{hh} + \mathcal{L}_{hhh} + \mathcal{L}_{hhhh} + \dots$



- Classical physics:** we use the saddle-point approximation in path integrals.



- Enjoy the advantages of **quantum field theory methods** and **classical physics**
powerful and systematic & **purely classical at all steps (simplicity)**

Effective Field Theory

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The in-in effective action is obtained by performing a closed-time-path integral

$$e^{i\mathcal{S}_{\text{eff}}[x_{a,1}, x_{a,2}]} = \int \mathcal{D}h_1 \mathcal{D}h_2 e^{i(S_{\text{GR}}[h_1] - S_{\text{GR}}[h_2] + S_{\text{WL}}[h_1, x_{a,1}] - S_{\text{WL}}[h_2, x_{a,2}])}$$

It is convenient to use the Keldysh basis

Galley PRL 110 (2013) 174301

$$\begin{aligned} h_{\mu\nu}^- &= \frac{1}{2}(h_{1\mu\nu} + h_{2\mu\nu}) & x_{a,+}^\alpha &= \frac{1}{2}(x_{a,1}^\alpha + x_{a,2}^\alpha) \\ h_{\mu\nu}^+ &= h_{1\mu\nu} - h_{2\mu\nu} & x_{a,-}^\alpha &= x_{a,1}^\alpha - x_{a,2}^\alpha \end{aligned}$$

for which the matrix of (classical) propagators for gravitons becomes

$$i \begin{pmatrix} 0 & -\Delta_{\text{adv}}(x-y) \\ -\Delta_{\text{ret}}(x-y) & 0 \end{pmatrix}$$

The worldline equations of motion:

Kälin-Neef-Porto JHEP 01 (2023) 140

$$m_i \ddot{x}_i^\mu(\tau) = -\eta^{\mu\nu} \frac{\delta S_{\text{eff, int}}[x_{a,\pm}]}{\delta x_{i,-}^\nu(\tau)} \Big|_{\text{PL}} \quad \Delta p_i^\mu = -\eta^{\mu\nu} \int_{-\infty}^{\infty} d\tau \frac{\delta S_{\text{eff, int}}[x_{a,\pm}]}{\delta x_{i,-}^\nu(\tau)} \Big|_{\text{PL}}$$

Physical Limit (PL): $x_{a,-} \rightarrow 0$, $x_{a,+} \rightarrow x_a$.

Closed-time path integrals

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EQUILIBRIUM AND NONEQUILIBRIUM FORMALISMS MADE UNIFIED

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Physics Reports

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周光召



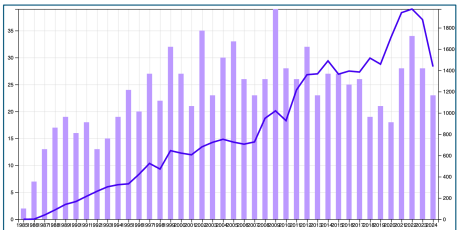
郝柏林



苏肇冰



于渌



Effective Field Theory

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- Effective action:



$$\Rightarrow -4\pi G m_1 m_2 \int d\tau_1 d\tau_2 \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot (x_1 - x_2)}}{k^2} \left(2(\dot{x}_1 \cdot \dot{x}_2)^2 - \dot{x}_1^2 \dot{x}_2^2 \right)$$

- Equation of motion:

$$x_i^\mu = b_i^\mu + u_i^\mu \tau_i + \delta x_i^\mu(\tau_i) + \dots$$

$$\begin{aligned} \ddot{x}_1^\mu &= 4i\pi G m_2 \int d\tau_2 \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot (b + u_1 \tau_1 + u_2 \tau_2)}}{k^2} \left[(2\gamma^2 - 1)k^\mu - 2(k \cdot u_1)(2\gamma u_2^\mu - u_1^\mu) \right] + \mathcal{O}(G^2) \\ &= 4i\pi G m_2 \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot b + i(k \cdot u_1)\tau_1} \hat{\delta}(k \cdot u_2)}{k^2} \left[(2\gamma^2 - 1)k^\mu - 2(k \cdot u_1)(2\gamma u_2^\mu - u_1^\mu) \right] + \mathcal{O}(G^2) \end{aligned}$$

- Trajectory to $\mathcal{O}(G)$:

$$\gamma = u_1 \cdot u_2, \quad u_i^2 = 1, \quad q \cdot u_i = 0$$

$$\begin{aligned} x_1^\mu &= b_1^\mu + u_1^\mu \tau_1 \\ &- 4i\pi G m_2 \left[(2\gamma^2 - 1)\eta^{\mu\nu} - 2(2\gamma u_2^\mu - u_1^\mu)u_1^\nu \right] \int \frac{d^D k}{(2\pi)^D} \frac{\hat{\delta}(k \cdot u_2)k_\nu}{k^2} \frac{e^{ik \cdot b + i(k \cdot u_1 - i0)\tau_1}}{(k \cdot u_1 - i0)^2} + \mathcal{O}(G^2) \end{aligned}$$

Effective Field Theory

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- Effective action:



$$\Rightarrow -4\pi G m_1 m_2 \int d\tau_1 d\tau_2 \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot (x_1 - x_2)}}{k^2} \left(2(\dot{x}_1 \cdot \dot{x}_2)^2 - \dot{x}_1^2 \dot{x}_2^2 \right)$$

- Impulse:

$$\Delta p_1^\mu = -\eta^{\mu\nu} \frac{\partial S}{\partial x_1^\nu} = 4\pi G m_1 m_2 \int d\tau_1 d\tau_2 \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot (x_1 - x_2)}}{k^2} \left(2(\dot{x}_1 \cdot \dot{x}_2)^2 - \dot{x}_1^2 \dot{x}_2^2 \right)$$

- LO (1PM):

$$x_i^\mu = b_i^\mu + u_i^\mu \tau_i + \delta x_i^\mu(\tau_i) + \dots$$

$$\Delta p_1^\mu = 4\pi G m_1 m_2 (2\gamma^2 - 1) \int \frac{d^D q}{(2\pi)^D} \frac{i q^\mu \hat{\delta}(q \cdot u_1) \hat{\delta}(q \cdot u_2) e^{iq \cdot b}}{q^2} = \frac{(2\gamma^2 - 1)}{\sqrt{\gamma^2 - 1}} \frac{2G m_1 m_2 b^\mu}{b^2}$$

- NLO (2PM):

$$\gamma = u_1 \cdot u_2, u_i^2 = 1, q \cdot u_i = 0$$

$$\Delta p_1^\mu \sim G^2 \int \frac{d^D q}{(2\pi)^D} \delta(q \cdot u_1) \delta(q \cdot u_2) e^{iq \cdot b} \int \frac{d^D k}{(2\pi)^D} \frac{\delta(k \cdot u_2)}{(k \cdot u_1 - i0)^\# [k^2] [(k - q)^2]}$$

Effective Field Theory

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- Effective action:



$$\Rightarrow -4\pi G m_1 m_2 \int d\tau_1 d\tau_2 \int \frac{d^D k}{(2\pi)^D} \frac{e^{ik \cdot (x_1 - x_2)}}{k^2} \left(2(\dot{x}_1 \cdot \dot{x}_2)^2 - \dot{x}_1^2 \dot{x}_2^2 \right)$$

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- LO (1PM):

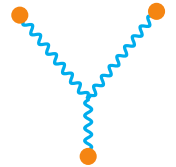
$$x_i^\mu = b_i^\mu + u_i^\mu \tau_i + \delta x_i^\mu(\tau_i) + \dots$$

$$\Delta p_1^\mu = 4\pi G m_1 m_2 (2\gamma^2 - 1) \int \frac{d^D q}{(2\pi)^D} \frac{i q^\mu \hat{\delta}(q \cdot u_1) \hat{\delta}(q \cdot u_2) e^{iq \cdot b}}{q^2} = \frac{(2\gamma^2 - 1)}{\sqrt{\gamma^2 - 1}} \frac{2G m_1 m_2 b^\mu}{b^2}$$

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Effective Field Theory

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- Observables at $\mathcal{O}(G^N)$

Kälin-ZL-Porto PRL2020

Dlapa-Kälin-ZL-Porto PRL2022 JHEP2024

$$\Delta p_i^\mu \sim \int d^D q \frac{e^{iq \cdot b} \delta(q \cdot u_1) \delta(q \cdot u_2)}{|q^2|^\sharp} \int \left(\prod_{i=1}^{N-1} d^D \ell_i \frac{\delta(\ell_j \cdot u_a)}{(\ell_j \cdot u_b - i0)^{\nu_i}} \right) \frac{\mathcal{N}^\mu(q, u_a)}{D_1 D_2 D_3 \cdots}$$

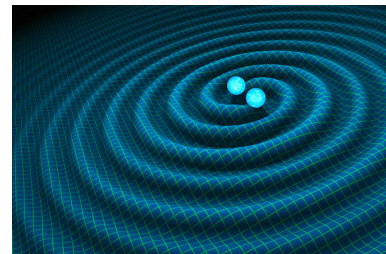
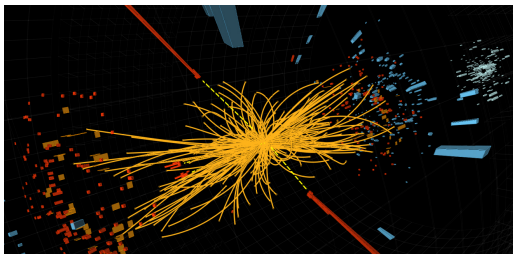
- ▶ Graviton propagators:

$$\frac{1}{D_i} \longrightarrow \frac{1}{\ell^2 + i0} \quad \text{or} \quad \frac{1}{(\ell^0 \pm i0)^2 - \vec{\ell}^2}$$

- ▶ **Cut**: always one delta function $\delta(\ell_i \cdot u_a)$ for each loop

- ▶ Kinematics: $q \cdot u_a = 0$, $u_a^2 = 1$, $u_1 \cdot u_2 = \gamma \implies$ **single scale** γ to all orders

- Multi-loop technology from QFT can be used to solve classical gravitational problems!

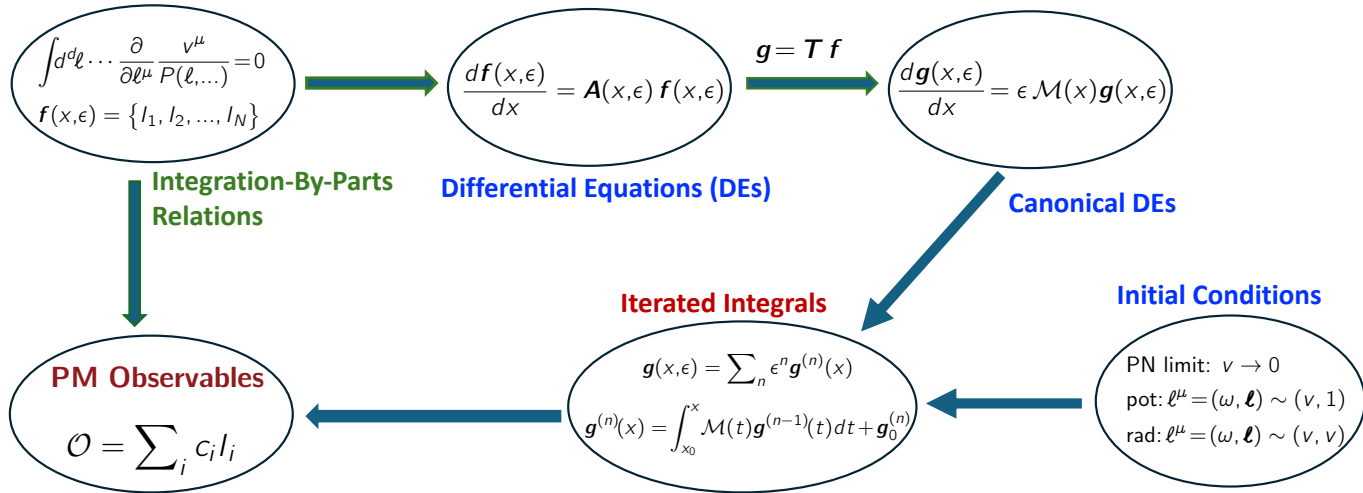


- Observables at $\mathcal{O}(G^N)$ Kälin-ZL-Porto PRL 2020 Dlapa-Kälin-ZL-Porto PRL 2022 Dlapa-Kälin-ZL-Neef-Porto PRL 2023

$$\Delta p_i^\mu \sim \int d^D q \frac{e^{iq \cdot b} \delta(q \cdot u_1) \delta(q \cdot u_2)}{|q^2|^\#} \int \left(\prod_{i=1}^{N-1} d^D \ell_i \frac{\delta(\ell_j \cdot u_a)}{(\ell_i \cdot u_b - i0)^{\nu_i}} \right) \frac{\mathcal{N}^\mu(q, u_a)}{D_1 D_2 D_3 \dots}$$

- Perturbative QFT toolbox:

Dlapa-Kälin-ZL-Porto JHEP 2024



- Post-Minkowskian physics can be bootstrapped from Post-Newtonian data using DEs!

Spin interactions

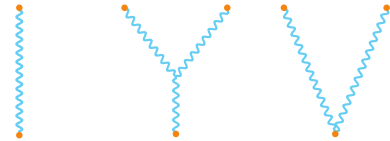
$$g_{\mu\nu} e_a^\mu e_b^\nu = \eta_{ab}$$

$$-\frac{1}{2} \left(\omega_\mu^{ab} S_{ab} v^\mu + \frac{1}{m} R_{\beta\rho\mu\nu} e_a^\alpha e_b^\beta e_c^\mu e_d^\nu S^{ab} S^{cd} v^\rho v_\alpha - \frac{C_{ES}}{m} E_{\mu\nu} e_a^\mu e_b^\nu S^{ac} S_c{}^b + \dots \right)$$

- EFT provides a systematic way to include spin effects.
- Needed one-loop integrals are simple

See Fei's talk

$$\int d^D \ell \frac{\delta(\ell \cdot u_1)}{[(\pm \ell \cdot u_2)]^{a_1}} \frac{1}{[\ell^2]^{a_2} [(\ell - q)^2]^{a_3}}$$



- Nontrivial to simplify complicated tensor expressions

Observables:

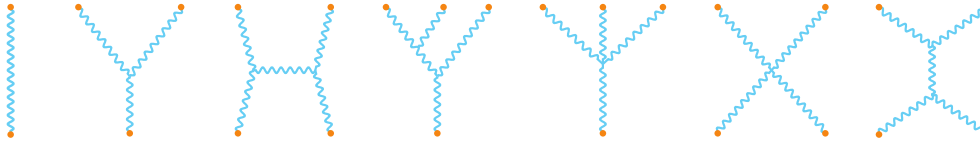
ZL-Porto-Yang JHEP 2021

$$\Delta p_1^\mu = \frac{\nu G^2 M^3}{|b|^3} \left[3D_1 \epsilon_{\alpha\rho\beta\sigma} \hat{b}^\mu \hat{b}^\alpha u_1^\beta u_2^\sigma a_1^\rho + \dots + \frac{D_{20}}{|b|} u_1^\mu (a_1 \cdot a_2) + \dots + \frac{D_{14}}{|b|} u_2^\mu a_1^2 + \dots \right]$$

$$s_\mu = m a_\mu \equiv \frac{1}{2m} \epsilon_{\mu\nu\alpha\beta} p^\nu S^{\alpha\beta}$$

NNLO: 3PM

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- $\mathcal{O}(G^3)$: two-loop integrals

Kälin-ZL-Porto PRL 2020 PRD 2020

$$\int \frac{d^D \ell_1 d^D \ell_2 \delta(\ell_1 \cdot u_1) \delta(\ell_2 \cdot u_2)}{[\ell_1 \cdot u_2]^{a_1} [\pm \ell_2 \cdot u_1]^{a_2}} \frac{1}{[\ell_1^2]^{a_3} [\ell_2^2]^{a_4} [(\ell_1 + \ell_2 - q)^2]^{a_5} [(\ell_1 - q)^2]^{a_6} [(\ell_2 - q)^2]^{a_7}}$$

- The reduction and evaluation of integrals can be performed in standard techniques.

$$\partial_x \vec{f}(x, \epsilon) = \epsilon \left(\frac{A}{x} + \frac{B}{x+1} + \frac{C}{x-1} \right) \vec{f}(x, \epsilon)$$

- Conservative dynamics at $\mathcal{O}(G^3)$:

Kälin-ZL-Porto PRL 2020

$$\Delta p_1^\mu = \frac{G^3 b^\mu}{|b|^2} \left(\frac{8m_1^2 m_2^2 (4\gamma^4 - 12\gamma^2 - 3)}{(\gamma^2 - 1)} \log(\gamma - \sqrt{\gamma^2 - 1}) - \frac{2m_1 m_2 (m_1^2 + m_2^2) (16\gamma^6 - 32\gamma^4 + 16\gamma^2 - 1)}{(\gamma^2 - 1)^{5/2}} \right. \\ \left. - \frac{4m_1^2 m_2^2 \gamma (20\gamma^6 - 90\gamma^4 + 120\gamma^2 - 53)}{3(\gamma^2 - 1)^{5/2}} \right) + \frac{3\pi (2\gamma^2 - 1) (5\gamma^2 - 1)}{2(\gamma^2 - 1)^2} \frac{G^3 m_1 m_2 (m_1 + m_2)}{|b|^2 |b|^{3/2}} \left((m_1 + \gamma m_2) u_2^\mu - (m_2 + \gamma m_1) u_1^\mu \right)$$

- We provided the first confirmation for the result from a scattering amplitude calculation.

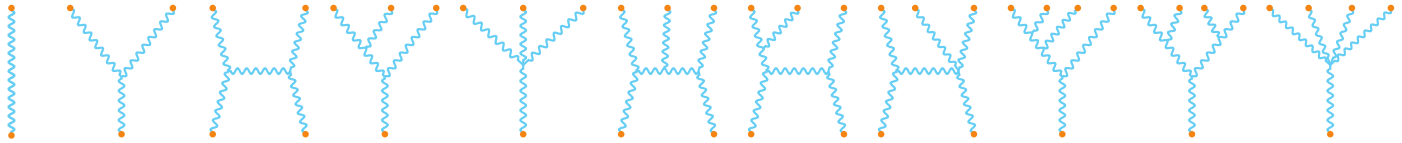
Bern-Cheung-Roiban-Shen-Solon-Zeng 2019

- We computed quadrupolar and octupole tidal corrections at $\mathcal{O}(G^3)$.

Kälin-ZL-Porto PRD 2020

NNNLO: 4PM

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$\mathcal{O}(G^4)$: three-loop integrals

Dlapa-Kälin-**ZL**-Porto PRL 2022 PLB 2022 Dlapa-Kälin-**ZL**-Neef-Porto PRL 2023

$$\int d^D \ell_1 d^D \ell_2 d^D \ell_3 \frac{\delta(\ell_1 \cdot u_1) \delta(\ell_2 \cdot u_2) \delta(\ell_3 \cdot u_2)}{[\ell_1 \cdot u_2]^{\alpha_1} [\ell_2 \cdot u_1]^{\alpha_2} [\ell_3 \cdot u_1]^{\alpha_3}} \frac{D_8^{-\nu_8} D_9^{-\nu_9}}{D_1^{\nu_1} D_2^{\nu_2} \dots D_7^{\nu_7}} \left\{ \ell_1^2, \ell_2^2, (\ell_1 - q)^2, (\ell_2 - q)^2, (\ell_3 - q)^2, \right. \\ \left. \ell_3^2, (\ell_1 - \ell_2)^2, (\ell_2 - \ell_3)^2, (\ell_3 - \ell_1)^2 \right\}$$

IBP reduction:

conservative: $\mathcal{O}(10^2)$ master integrals

full: $\mathcal{O}(10^3)$ master integrals

Differential Equations

Dlapa-Kälin-**ZL**-Porto PRL 2022 PLB 2022 Dlapa-Henn-Wagner 2211.16357

$$\frac{d\vec{f}(x, \epsilon)}{dx} = \epsilon \mathcal{M}(x) \vec{f}(x, \epsilon)$$

- The majority can be solved in terms of **multiple polylogarithms**.
- Elliptic integrals appear in post-Minkowskian gravity for the first time.

Elliptic differential equations

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- From 4PM order, more complicated functions appear beyond polylogarithms.
- An elliptic example:

$$\frac{d}{dx} \begin{pmatrix} f_1(x) \\ f_2(x) \\ f_3(x) \end{pmatrix} = \begin{pmatrix} \frac{1-x^2}{2x(1+x^2)} & \frac{1+x^2}{4x(1-x^2)} & \frac{3x}{(1-x^2)(1+x^2)} \\ -\frac{1-x^2}{x(1+x^2)} & \frac{3(1+x^2)}{2x(1-x^2)} & -\frac{6x}{(1-x^2)(1+x^2)} \\ \frac{1-x^2}{x(1+x^2)} & -\frac{1+x^2}{2x(1-x^2)} & -\frac{1-4x^2+x^4}{x(1-x^2)(1+x^2)} \end{pmatrix} \begin{pmatrix} f_1(x) \\ f_2(x) \\ f_3(x) \end{pmatrix} + \mathcal{O}(\epsilon)$$

It can then be written as a third-order differential equation:

$$\left[\frac{d^3}{dx^3} - \frac{6x}{1-x^2} \frac{d^2}{dx^2} - \frac{1-4x^2+7x^4}{x^2(1-x^2)^2} \frac{d}{dx} - \frac{1+x^2}{x^3(1-x^2)} \right] f_1(x) = 0$$

It is easy to find the three solutions:

$$x K^2(1-x^2), \quad x K(1-x^2)K(x^2), \quad x K^2(x^2)$$

Complete elliptic integrals: $K(x) \equiv \int_0^1 \frac{dt}{\sqrt{(1-t^2)(1-xt^2)}}$

Elliptic differential equations

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With the knowledge of leading- ϵ solutions, one may transform the elliptic diagonal block into

$$\frac{d}{dx} \vec{g}(x, \epsilon) = \epsilon \tilde{D}_{\text{ell}}(x) \vec{g}(x, \epsilon) + \dots$$

with

$$\tilde{D}_{\text{ell}} = \begin{pmatrix} -\frac{4(1+x^2)}{3x(1-x^2)} & \frac{\pi^2}{x(1-x^2)K^2(1-x^2)} & 0 \\ \frac{2(1+110x^2+x^4)K^2(1-x^2)}{3\pi^2x(1-x^2)} & -\frac{4(1+x^2)}{3x(1-x^2)} & \frac{\pi^2}{x(1-x^2)K^2(1-x^2)} \\ \frac{16(1+x^2)(1-18x+x^2)(1+18x+x^2)K^4(1-x^2)}{27\pi^2x(1-x^2)} & \frac{2(1+110x^2+x^4)K^2(1-x^2)}{3\pi^2x(1-x^2)} & -\frac{4(1+x^2)}{3x(1-x^2)} \end{pmatrix}$$

- Elliptic integrals appear in the transformation matrix: automated by INITIAL
- Higher $\mathcal{O}(\epsilon)$: Iterated integrals involving elliptic kernels.

See Xing's talk

NNNLO: 4PM

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The full impulse at $\mathcal{O}(G^4)$: Dlapa-Kälin-ZL-Porto PRL 2022 JHEP 2023 Dlapa-Kälin-ZL-Neef-Porto PRL 2023

$$\Delta p_1^\mu|_{\text{NNNLO}} = \frac{G^4}{|b|^4} \left(c_b \frac{b^\mu}{|b|} + c_1 \frac{\gamma u_2^\mu - u_1^\mu}{\gamma^2 - 1} + c_2 \frac{\gamma u_1^\mu - u_2^\mu}{\gamma^2 - 1} \right)$$

$$\begin{aligned} \frac{c_b}{\pi} = & -\frac{3h_{34}m_2m_1(m_1^3+m_2^3)}{64v_\infty^5} + \frac{m_1^2m_{12}m_2^2}{4} \left[\frac{3h_6K^2(w_2)}{4v_\infty^3} - \frac{3h_8K(w_2)E(w_2)}{4v_\infty^3} + \frac{21h_5w_3E^2(w_2)}{8v_\infty^3} - \frac{\pi^2h_{16}v_\infty}{4(\gamma+1)} + \frac{3\gamma h_{10}(Li_2(w_2) - 4Li_2(\sqrt{w_2}))}{w_3v_\infty^2} \right. \\ & + \log(v_\infty) \left(\frac{h_{32}}{2v_\infty^3} - \frac{3h_{14}\log(\frac{w_3}{2})}{v_\infty} - \frac{3\gamma h_{22}\log(w_1)}{2v_\infty^4} \right) \Bigg] + m_2^2m_1^3 \left[\frac{h_{52}}{48v_\infty^6} - \frac{h_{63}}{768\gamma^9w_3v_\infty^5} - \frac{3v_\infty(h_{40}Li_2(w_2) + 2w_3h_{33}Li_2(-w_2))}{64w_3} \right. \\ & + \frac{3h_{14}\log(\frac{w_3}{2})\log(w_3)}{4v_\infty} + \frac{\gamma h_{39}\log(w_1)}{8w_3^3v_\infty^2} + \frac{3\gamma h_{22}\log(w_3)\log(w_1)}{8v_\infty^4} - \frac{h_{35}\log(\frac{w_3}{2})}{32v_\infty^5} + \frac{h_{56}\log(2) - h_{57}\log(w_3) + 2\gamma h_{55}\log(\gamma)}{32v_\infty^5} - \frac{\gamma h_{51}\log(w_1)}{16v_\infty^7} \Bigg] \\ & + m_1^2m_2^3 \left[\frac{h_{58}}{192\gamma^7v_\infty^5} + \frac{h_{53}}{48v_\infty^6} + \frac{\gamma h_{49}\log(w_1)}{16v_\infty^6} - \frac{2\gamma h_{50}\log(w_1) + 3\gamma^2h_{13}\log^2(w_1)}{32v_\infty^7} - \frac{h_{41}\log(\frac{w_3}{2})}{8v_\infty^4} + \frac{3\gamma\log(w_1)(5h_{26}\log(2) + 8h_{12}\log(w_3))}{8v_\infty^4} \right. \\ & - \frac{h_{36}\log(w_3)}{4v_\infty^3} + \frac{\gamma h_{30}\log(\gamma)}{2v_\infty^3} + \frac{h_{37}\log(2)}{8v_\infty^3} + \frac{3(h_{17}w_3Li_2(w_2) - 2h_{23}Li_2(-w_2) + h_{15}\log^2(w_3) - h_9\log^2(2))}{8v_\infty} - \frac{3h_7\log(2)\log(w_3)}{v_\infty} \Bigg] \\ c_1 = & m_1m_2^2 \left(\frac{2h_{46}m_{12s}}{v_\infty^6} + \frac{9\pi^2h_1m_{12}^2}{32v_\infty^2} \right) + m_1^2m_2^3 \left(\frac{4\gamma h_{47}}{3v_\infty^6} - \frac{8\gamma h_2\log(w_1)}{v_\infty^6} + \frac{16h_{25}\log(w_1)}{v_\infty^3} - \frac{8h_3}{3v_\infty^5} \right) \\ c_2 = & -m_1^4m_2 \left(\frac{9\pi^2h_1}{32v_\infty^2} + \frac{2h_{46}}{v_\infty^6} \right) + m_2^2m_1^3 \left[\frac{h_{60}}{705600\gamma^8v_\infty^5} - \frac{4\gamma h_{48}}{3v_\infty^6} + \frac{3h_{38}(Li_2(w_2) - 4Li_2(\sqrt{w_2})) - \gamma h_{21}(Li_2(-w_1^2) + 2\log(\gamma)\log(w_1))}{16v_\infty^4} \right. \\ & + \frac{3\gamma h_{31}(2Li_2(-w_1) + \log(w_1)\log(w_3))}{8v_\infty^4} + \frac{h_{62}\log(w_1)}{6720\gamma^9v_\infty^6} + \frac{32\gamma^2h_{44}\log^2(w_1)}{v_\infty^7} + \frac{8\gamma(2h_4\log(2) - h_{27}\log(w_1))\log(w_1)}{v_\infty^4} - \frac{32h_{29}\log(w_1)}{3v_\infty^3} + \frac{\pi^2h_{42}}{192v_\infty^4} \Bigg] \\ & + m_2^3m_1^2 \left[\frac{h_{59}}{1440\gamma^7v_\infty^5} - \frac{h_{19}(Li_2(-w_1^2) + 2\log(\gamma)\log(w_1))}{8v_\infty^4} + \frac{h_{43}(Li_2(w_2) - 4Li_2(\sqrt{w_2}))}{32v_\infty^4} - \frac{h_{20}(2Li_2(-w_1) + \log(w_1)\log(w_3))}{4v_\infty^4} \right. \\ & - \frac{h_{61}\log(w_1)}{480\gamma^8v_\infty^6} - \frac{16\gamma^2h_{11}\log^2(w_1)}{v_\infty^4} - \frac{32\gamma h_{45}\log^2(w_1)}{v_\infty^7} + \frac{16\gamma h_{28}\log(w_1)}{5v_\infty^3} - \frac{32h_{24}\log(2)\log(w_1)}{v_\infty^4} - \frac{\pi^2h_{18}}{48v_\infty^4} - \frac{2h_{54}}{45v_\infty^6} \Bigg] \end{aligned}$$

with $\gamma \equiv u_1 \cdot u_2$, $v_\infty = \sqrt{\gamma^2 - 1}$, $w_1 = \gamma - v_\infty$, $w_2 = \frac{\gamma-1}{\gamma+1}$, $w_3 = \gamma + 1$, h_i = polynomial in γ .

$$Li_2(z) \equiv \int_0^1 \frac{dx \log(1-x)}{1-x}$$

$$K(z) \equiv \int_1^\infty \frac{dx}{\sqrt{(1-x^2)(1-zx^2)}}$$

$$E(k) \equiv \int_0^1 dx \frac{\sqrt{1-zx^2}}{\sqrt{1-x^2}}$$

NNNLO: 4PM

止於至善

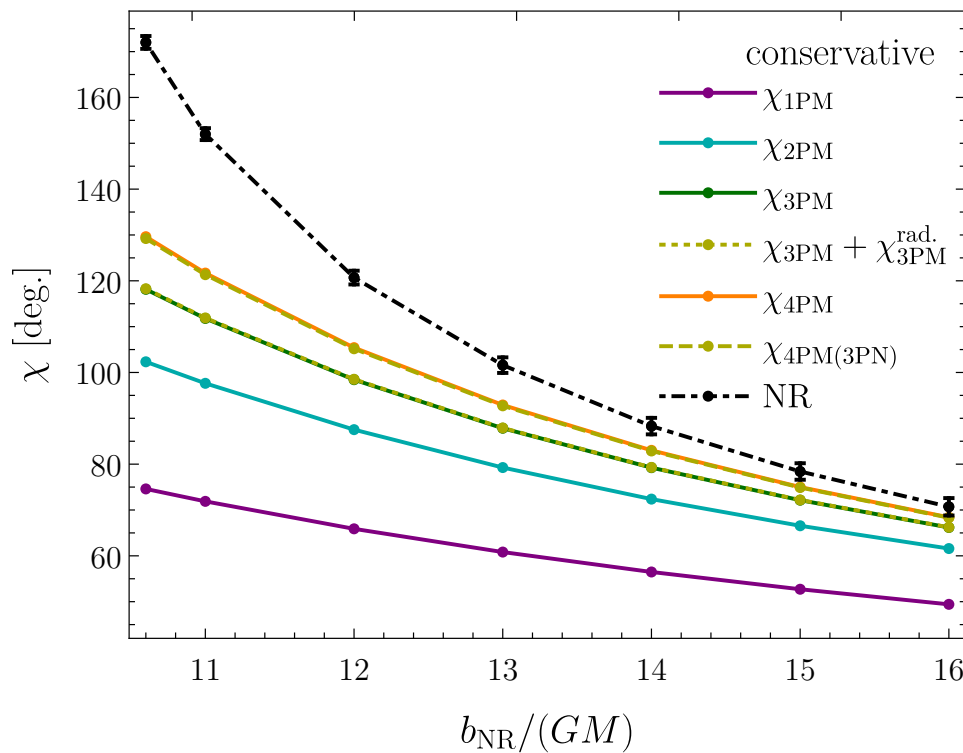
The full impulse at $\mathcal{O}(G^4)$: [Dlapa-Kälin-ZL-Porto JHEP 2023](#) [Dlapa-Kälin-ZL-Neef-Porto PRL 2023](#)

$$\Delta p_1^\mu|_{\text{NNNLO}} = \frac{G^4}{|b|^4} \left(c_b \frac{b^\mu}{|b|} + c_1 \frac{\gamma u_2^\mu - u_1^\mu}{\gamma^2 - 1} + c_2 \frac{\gamma u_1^\mu - u_2^\mu}{\gamma^2 - 1} \right)$$

- We obtained the full dynamics of binary inspirals at $\mathcal{O}(G^4)$ for the first time.
- Conservative part agrees perfectly with previous derivations.
[Bern-Parra-Martinez-Roiban-Ruf-Shen-Solon-Zeng 2022](#) [Dlapa-Kälin-ZL-Porto PRL 2022](#)
- Perfect agreement with the state-of-the-art PN computations
[Cho-Dandapat-Gopakumar 2021](#) [Cho 2022](#) [Bini-Geralico 2021 2022](#) [Bini-Damour 2022](#)
- Later, two new calculations confirmed our results.
[Damgaard-Hansen-Planté-Vanhove 2023](#) (exponentiation of amplitudes)
[Jakobsen-Mogull-Plefka-Sauer-Xu 2023](#) (worldline formalism)

Analytic vs Numerical Relativity

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Khalil-Buonanno-Steinhoff-Vines 2204.05047

NNNLO: local-in-time part

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- The full result does not describe generic elliptic-like motion due to nonlocal-in-time effects.

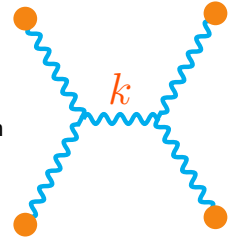
Damour-Jaranowski-Schäfer 2014 Galley-Leibovich-Porto-Ross 2015 Cho-Kälin-Porto 2021

- Nonlocal-in-time radial action:

$$\mathcal{S}_r^{(\text{nloc})} = -\frac{GE}{2\pi} \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{dE}{d\omega} \log \left(\frac{4\omega^2}{\mu^2} e^{2\gamma_E} \right)$$

- The 4PM integrand can be built from 3PM diagrams.

$$\omega \equiv k \cdot u_{\text{com}}$$



$$\int d^D \ell_1 d^D \ell_2 \frac{\delta(\ell_1 \cdot u_1) \delta(\ell_2 \cdot u_2)}{[\ell_1 \cdot u_2][\ell_2 \cdot u_1]} \frac{\log(\omega^2)}{[\ell_1^2][\ell_2^2][(\ell_1 + \ell_2 - q)^2][(\ell_1 - q)^2][(\ell_2 - q)^2]}$$

- We managed to compute the integrals and obtained nonlocal-in-time contribution:

$$\begin{aligned} & \frac{\nu}{(\gamma^2 - 1)^2} \left[h_1 + \frac{\pi^2 h_2}{\sqrt{\gamma^2 - 1}} + h_3 \log \frac{\gamma + 1}{2} + \frac{h_4 \text{arccosh}(\gamma)}{\sqrt{\gamma^2 - 1}} + h_5 \log \frac{\gamma - 1}{8} + h_6 \log^2 \frac{\gamma + 1}{2} + \frac{h_8 \log(2) \text{arccosh}(\gamma)}{\sqrt{\gamma^2 - 1}} \right. \\ & \left. + h_7 \text{arccosh}(\gamma)^2 + h_9 \log \frac{\gamma - 1}{8} \log \frac{\gamma + 1}{2} + \frac{h_{10} \log \frac{\gamma^2 - 1}{16} \text{arccosh}(\gamma)}{\sqrt{\gamma^2 - 1}} + h_{11} \text{Li}_2 \frac{\gamma - 1}{\gamma + 1} + h_{12} \frac{\text{arccosh}^2(\gamma) + 4 \text{Li}_2(\sqrt{\gamma^2 - 1} - \gamma)}{\sqrt{\gamma^2 - 1}} \right] \end{aligned}$$

Coefficients h_i : exact- ν (iterated elliptic integrals) and SF-expanded (30SF) forms

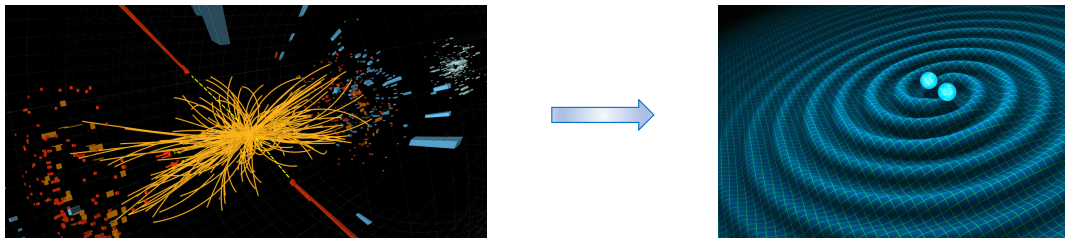
Dlapa-Kälin-ZL-Porto PRL 2024

- Using 6PN results in the literature, we constructed an improved bound Hamiltonian.

Conclusion & Outlook

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Modern techniques from Quantum Field Theory have already proven useful to improve theoretical predictions for gravitational-wave observables.



We have developed an efficient framework and made breakthroughs to NNNLO.

- Conservative spin & tidal effects at NLO [JHEP 06 \(2021\) 012](#) [PRD 102 \(2020\) 124025](#)
- Conservative dynamics at NNLO [PRL 125 \(2020\) 261103](#)
- Conservative dynamics at NNNLO [PLB 822 \(2021\) 136698](#) [PRL 128 \(2022\) 161104](#) [PRL 130 \(2023\) 101401](#)
- Local-in-time & nonlocal-in-time separation to $N^4\text{LO}$ [PRL 132 \(2024\) 221401](#) [arXiv:2506.20665](#)
- Novel techniques to evaluate classical loop integrals [JHEP 07 \(2023\) 181](#) [JHEP 08 \(2023\) 109](#)

Outlook: precision frontier

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GW science is just starting! discovery potential = precise theoretical predictions!

	0PN	1PN	2PN	3PN	4PN	5PN	6PN	...
1PM	$(\frac{GM}{r}) \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots)$							
2PM		$(\frac{GM}{r})^2 \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots)$						
3PM			$(\frac{GM}{r})^3 \times (1 + v^2 + v^4 + v^6 + v^8 + \dots)$					
4PM				$(\frac{GM}{r})^4 \times (1 + v^2 + v^4 + v^6 + \dots)$				
5PM					$(\frac{GM}{r})^5 \times (1 + v^2 + v^4 + \dots)$			
6PM						$(\frac{GM}{r})^6 \times (1 + v^2 + \dots)$		
7PM							$(\frac{GM}{r})^7 \times (1 + \dots)$	

0SF 1SF 2SF 3SF ...

► 1SF: [Driesse-Jakobsen-Mogull-Plefka-Sauer-Usovitsch 2024](#) [D-J-Klemm-M-Nega-P-S-U 2024](#)

Outlook: precision frontier

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3PM	$(\frac{GM}{r})^3 \times (1 + v + v^2 + v^3 + v^4 + v^5 + \dots)$							
4PM	$(\frac{GM}{r})^4 \times (1 + v + v^2 + v^3 + v^4 + v^5 + \dots)$							
5PM	$(\frac{GM}{r})^5 \times (1 + v + v^2 + v^3 + v^4 + v^5 + \dots)$							
6PM	$(\frac{GM}{r})^6 \times (1 + v + v^2 + v^3 + v^4 + v^5 + \dots)$							
7PM	$(\frac{GM}{r})^7 \times (1 + v + v^2 + v^3 + v^4 + v^5 + \dots)$							

0SF 1SF 2SF 3SF ...



► 1SF: Driesse-Jakobsen-Mogull-Plefka-Sauer-Usovitsch 2024

D-J-Klemm-M-Nega-P-S-U 2024

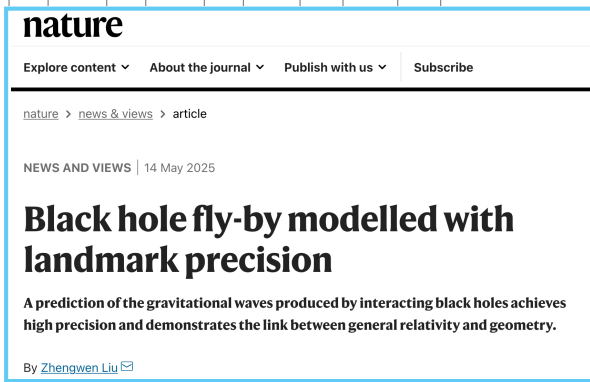
Outlook: precision frontier

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GW science is just starting! discovery potential = precise theoretical predictions!

	0PN	1PN	2PN	3PN	4PN	5PN	6PN	...	
1PM		$\left(\frac{GM}{r}\right) \times (1$	$+ v^2$	$+ v^4$	$+ v^6$	$+ v^8$	$+ v^{10}$	$+ v^{12} + \dots)$	$= \left(\frac{GM}{r}\right) \times (1)$
2PM		$\left(\frac{GM}{r}\right)^2 \times (1$	$+ v^2$	$+ v^4$	$+ v^6$	$+ v^8$	$+ v^{10} + \dots)$	$= \left(\frac{GM}{r}\right)^2 \times (1)$	
3PM									$\left(\frac{GM}{r}\right)^3 \times (1 + \nu)$
4PM									$\left(\frac{GM}{r}\right)^4 \times (1 + \nu)$
5PM									$\left(\frac{GM}{r}\right)^5 \times (1 + \nu + \nu^2)$
6PM									$\left(\frac{GM}{r}\right)^6 \times (1 + \nu + \nu^2)$
7PM									$\left(\frac{GM}{r}\right)^7 \times (1 + \nu + \nu^2 + \nu^3)$

0SF 1SF 2SF 3SF ...



- 1SF: Driesse-Jakobsen-Mogull-Plefka-Sauer-Usovitsch 2024 D-J-Klemm-M-Nega-P-S-U 2024

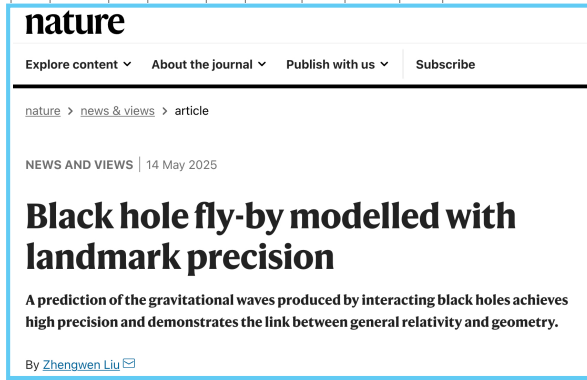
Outlook: precision frontier

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GW science is just starting! discovery potential = precise theoretical predictions!

$$\begin{array}{l}
 \text{0PN} \quad \text{1PN} \quad \text{2PN} \quad \text{3PN} \quad \text{4PN} \quad \text{5PN} \quad \text{6PN} \quad \dots \\
 \text{1PM} \quad \left(\frac{GM}{r}\right) \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots) = \left(\frac{GM}{r}\right) \times (1) \\
 \text{2PM} \quad \left(\frac{GM}{r}\right)^2 \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) = \left(\frac{GM}{r}\right)^2 \times (1) \\
 \text{3PM} \quad \left(\frac{GM}{r}\right)^3 \times (1 + v) \\
 \text{4PM} \quad \left(\frac{GM}{r}\right)^4 \times (1 + v) \\
 \text{5PM} \quad \left(\frac{GM}{r}\right)^5 \times (1 + v + v^2) \\
 \text{6PM} \quad \left(\frac{GM}{r}\right)^6 \times (1 + v + v^2) \\
 \text{7PM} \quad \left(\frac{GM}{r}\right)^7 \times (1 + v + v^2 + v^3)
 \end{array}$$

0SF 1SF 2SF 3SF ...



- ▶ 1SF: Driesse-Jakobsen-Mogull-Plefka-Sauer-Usovitsch 2024 D-J-Klemm-M-Nega-P-S-U 2024
- ▶ 1SF: local-in-time conservative dynamics: Dlapa-Kälin-ZL-Porto 2506.20665

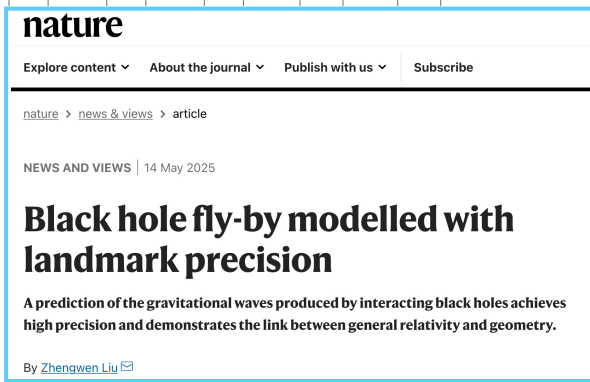
Outlook: precision frontier

心於至善

GW science is just starting! discovery potential = precise theoretical predictions!

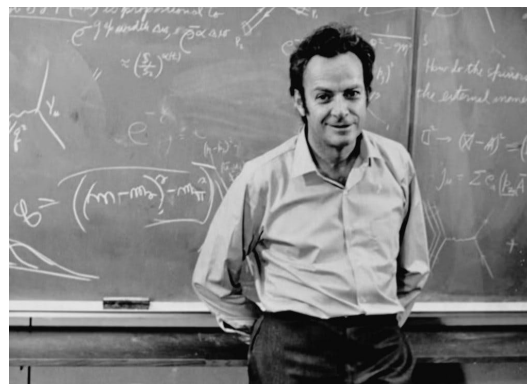
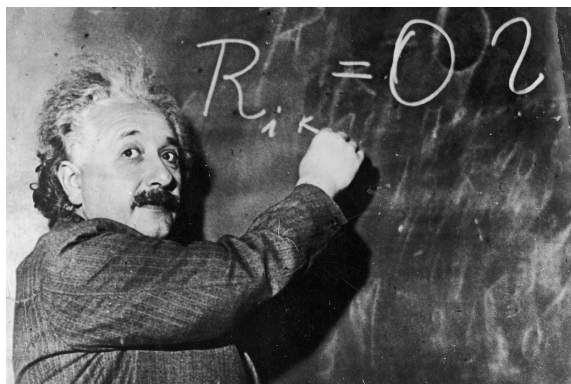
$$\begin{array}{l}
 0\text{PN} \quad 1\text{PN} \quad 2\text{PN} \quad 3\text{PN} \quad 4\text{PN} \quad 5\text{PN} \quad 6\text{PN} \quad \dots \\
 1\text{PM} \quad \left(\frac{GM}{r}\right) \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + v^{12} + \dots) = \left(\frac{GM}{r}\right) \times (1) \\
 2\text{PM} \quad \left(\frac{GM}{r}\right)^2 \times (1 + v^2 + v^4 + v^6 + v^8 + v^{10} + \dots) = \left(\frac{GM}{r}\right)^2 \times (1) \\
 3\text{PM} \quad \left(\frac{GM}{r}\right)^3 \times (1 + v) \\
 4\text{PM} \quad \left(\frac{GM}{r}\right)^4 \times (1 + v) \\
 5\text{PM} \quad \left(\frac{GM}{r}\right)^5 \times (1 + v + v^2) \\
 6\text{PM} \quad \left(\frac{GM}{r}\right)^6 \times (1 + v + v^2) \\
 7\text{PM} \quad \left(\frac{GM}{r}\right)^7 \times (1 + v + v^2 + v^3)
 \end{array}$$

0SF 1SF 2SF 3SF ...



- ▶ 1SF: Driesse-Jakobsen-Mogull-Plefka-Sauer-Usovitsch 2024 D-J-Klemm-M-Nega-P-S-U 2024
- ▶ 1SF: local-in-time conservative dynamics: Dlapa-Kälin-ZL-Porto 2506.20665
- ▶ 2SF: nonplanar diagram & more complicated functions

Feynman's integrals solve Einstein's equations!



谢谢!