# A new channel to search for dark matter at Belle II







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in collaboration with Jinhan Liang and Lan Yang [2212.04252]

### Sensitivity on invisible dark photon models





# 1 New dark matter channel



Most studies focus on mono-X channel with SM X produced at the primary vertex









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SM







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Different mono-X channels

- mono-photon
- mono-jet
- mono-Higgs
- mono-Z
- mono-top













SM





SM

A pair of SM particles produced at the primary vertex



SM



One SM particle interacts with the detector to produce a pair of DM particles

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### fixed target in collider



































• 
$$e^+e^- \rightarrow e^+e^-$$

• *e*<sup>-</sup> deposit energy in ECL







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### disappearing positron track

















#### • *e*<sup>-</sup>: CDC & ECL









#### • *e*<sup>-</sup>: CDC & ECL

#### • $e^+$ : CDC & ECL









#### • *e*<sup>-</sup>: CDC & ECL

#### • $e^+$ : CDC & ECL

### CDC: $\frac{\delta p_T}{dm} \simeq 0.4\%$ for $p_T \simeq 3$ GeV $p_T$









- *e*<sup>-</sup>: CDC & ECL
- $e^+$ : CDC & ECL

$$\text{CDC:} \frac{\delta p_T}{p_T} \simeq 0.4 \ \% \ \text{for} \ p_T \simeq 3 \ \text{GeV}$$

Equal & opposite momenta for  $e^-$  &  $e^+$  in the CM frame









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$$\text{CDC:}\, \frac{\delta p_T}{p_T} \simeq 0.4\,\% \,\, \text{for}\, p_T \simeq 3 \,\, \text{GeV}$$

Equal & opposite momenta for  $e^-$  &  $e^+$  in the CM frame

• missing energy: <5%  $e^+$  energy in ECL







### **Positron interaction with ECL**



#### annihilation w/ atomic electrons



bremsstrahlung w/ target nucleus



### ECL barrel: $32.2^{\circ} < \theta < 128.7^{\circ}$





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Better hermiticity (non-projective gaps between ECL crystals)





### ECL barrel: $32.2^{\circ} < \theta < 128.7^{\circ}$



Less non-instrumented setups (e.g., magnetic wires) between ECL & KLM

Better hermiticity (non-projective gaps between ECL crystals)

More beam BG in Endcaps





### Bhabha scattering



### $6 \times 10^{11} e^+ e^-$ in the barrel region with 50/ab







# 2 Background







# BG: $e^+$ + ECL $\rightarrow$ SM which then escape detection







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Neutrino BG is negligible (xsec is small)





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Neutrino BG is negligible (xsec is small)

Main BG is due to  $n/\gamma$ 









#### Photon energy measured in ECL





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ECL = 16- $X_0$  Csl crystals, w/  $X_0$  = 1.86 cm

Photon can also be detected by KLM

KLM = alternating sandwich of 4.7-cm iron plates and active detectors





#### Photon escapes ECL





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Photon energy spectrum due to  $e^+$  collision with ECL [Tsai & Whitis 1966]

$$\frac{dN_{\gamma}}{dx_{\gamma}}(t, x_{\gamma}) \simeq \frac{1}{x_{\gamma}} \frac{(1 - x_{\gamma})^{(4/3)t} - e^{-(7/9)t}}{7/9 + (4/3)\ln(1 - x_{\gamma})}$$
$$x_{\gamma} = E_{\gamma}/E_{e} \qquad tX_{0} \text{ is the distance}$$





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 $x_{\gamma} = E_{\gamma}/E_e$   $tX_0$  is the distance

$$\int_{0.95}^{1} dx_{\gamma} \frac{dN_{\gamma}}{dx_{\gamma}} (t = 16, x_{\gamma}) \simeq 4.7 \times 10^{-8}$$

 $\sim 2.8 \times 10^4 \gamma$ -BG after ECL for  $6 \times 10^{11} e^+$ 



![](_page_46_Picture_7.jpeg)

![](_page_47_Figure_1.jpeg)

![](_page_47_Picture_2.jpeg)

![](_page_47_Picture_3.jpeg)

#### GeV $\gamma$ is unlikely to penetrate the KLM

![](_page_48_Figure_2.jpeg)

![](_page_48_Picture_3.jpeg)

![](_page_48_Picture_4.jpeg)

GeV  $\gamma$  is unlikely to penetrate the KLM

However,  $\gamma$  can be absorbed by noninstrumented setups (e.g., magnet coil)

![](_page_49_Figure_3.jpeg)

![](_page_49_Picture_4.jpeg)

![](_page_49_Picture_5.jpeg)

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KLM veto power is limited

![](_page_50_Figure_4.jpeg)

![](_page_50_Picture_5.jpeg)

![](_page_50_Picture_6.jpeg)

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KLM veto power is limited

IFR @ BaBar, veto eff =  $4.5 \times 10^{-4}$ 

![](_page_51_Figure_5.jpeg)

![](_page_51_Picture_6.jpeg)

![](_page_51_Picture_7.jpeg)

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However,  $\gamma$  can be absorbed by noninstrumented setups (e.g., magnet coil)

KLM veto power is limited

IFR @ BaBar, veto eff =  $4.5 \times 10^{-4}$ 

13 photon BG (conservative)

![](_page_52_Figure_6.jpeg)

![](_page_52_Picture_7.jpeg)

![](_page_52_Picture_8.jpeg)

![](_page_53_Figure_1.jpeg)

![](_page_53_Picture_2.jpeg)

![](_page_53_Picture_3.jpeg)

GEANT4 simulation of  $10^9 e^+$  with 4.35 GeV onto a CsI target with  $1 X_0$ 

![](_page_54_Figure_2.jpeg)

![](_page_54_Picture_3.jpeg)

![](_page_54_Picture_4.jpeg)

GEANT4 simulation of  $10^9 e^+$  with 4.35 GeV onto a CsI target with  $1 X_0$ 

Full simulation with 16  $X_0$  is time-consuming

![](_page_55_Figure_3.jpeg)

![](_page_55_Picture_4.jpeg)

GEANT4 simulation of  $10^9 e^+$  with 4.35 GeV onto a CsI target with  $1 X_0$ 

Full simulation with 16  $X_0$  is time-consuming

Neutrons with significant energy are likely to be produced in the 1st  $X_0$  (confirmed in simulations with 2- $X_0$ )

![](_page_56_Figure_4.jpeg)

![](_page_56_Picture_5.jpeg)

![](_page_56_Picture_6.jpeg)

![](_page_57_Figure_1.jpeg)

![](_page_57_Picture_2.jpeg)

![](_page_57_Picture_3.jpeg)

At least 1 neutron with energy > 3 GeV

![](_page_58_Figure_2.jpeg)

![](_page_58_Picture_3.jpeg)

![](_page_58_Picture_4.jpeg)

At least 1 neutron with energy > 3 GeV

Energy deposition in ECL < 5%

![](_page_59_Figure_3.jpeg)

![](_page_59_Picture_4.jpeg)

![](_page_59_Picture_5.jpeg)

At least 1 neutron with energy > 3 GeV

Energy deposition in ECL < 5%

Veto  $p/\pi^{\pm}$  with momentum > 0.6 GeV (either deposit energy in ECL or produce tracks in KLM)

n  $e^{\neg}$ CDC ECL KLM

![](_page_60_Picture_5.jpeg)

![](_page_60_Picture_6.jpeg)

At least 1 neutron with energy > 3 GeV

Energy deposition in ECL < 5%

Veto  $p/\pi^{\pm}$  with momentum > 0.6 GeV (either deposit energy in ECL or produce tracks in KLM)

Count # of neutrons with K.E. > 280 MeV (hadronic shower threshold)

![](_page_61_Figure_5.jpeg)

![](_page_61_Picture_6.jpeg)

![](_page_61_Picture_7.jpeg)

![](_page_62_Figure_1.jpeg)

![](_page_62_Picture_2.jpeg)

![](_page_62_Picture_3.jpeg)

Prob to penetrate a target with length L

$$P = \exp(-L/\lambda_0)$$

 $\lambda_0$  = hadronic interaction length

KLM has ~  $3.9 \lambda_0$ 

ECL has  $\sim 0.8 \lambda_0$ 

![](_page_63_Figure_6.jpeg)

![](_page_63_Picture_7.jpeg)

![](_page_63_Picture_8.jpeg)

Prob to penetrate a target with length L

$$P = \exp(-L/\lambda_0)$$

 $\lambda_0$  = hadronic interaction length

KLM has  $\sim 3.9 \lambda_0$ 

ECL has  $\sim 0.8 \lambda_0$ 

Prob to penetrate ECL & KLM is about 1%

![](_page_64_Figure_7.jpeg)

![](_page_64_Picture_8.jpeg)

![](_page_64_Picture_9.jpeg)

![](_page_64_Picture_10.jpeg)

Prob to penetrate a target with length L

$$P = \exp(-L/\lambda_0)$$

 $\lambda_0$  = hadronic interaction length

KLM has  $\sim 3.9 \lambda_0$ 

ECL has  $\sim 0.8 \lambda_0$ 

Prob to penetrate ECL & KLM is about 1%

about 81 neutron background in total

![](_page_65_Figure_8.jpeg)

![](_page_65_Picture_9.jpeg)

![](_page_65_Picture_10.jpeg)

![](_page_65_Picture_11.jpeg)

## Summary on background estimation

#### BG: $e^+$ + ECL $\rightarrow \gamma/n$ which escape detection

Use KLM to veto such BG

- photon BG events:  $\sim 13$
- neutron BG events:  $\sim 81$

#### [Liang, ZL, Yang, 2212.04252]

![](_page_66_Figure_6.jpeg)

![](_page_66_Picture_7.jpeg)

![](_page_67_Picture_0.jpeg)

## Sensitivity on invisible dark photon

![](_page_67_Picture_2.jpeg)

#### Invisible dark photon

 $\mathscr{L}_{\text{int}} = A'_{\mu}(eQ_f\epsilon \bar{f}\gamma^{\mu}f + g_{\chi}\bar{\chi}\gamma^{\mu}\chi)$ 

dark photon  $A'_{\mu}$ 

suppressed coupling  $\epsilon$  to SM fermion

gauge coupling to hidden fermion  $\chi: g_{\chi} \gg e\epsilon$ 

$$m_{A'} = 3m_{\chi}$$

![](_page_68_Picture_7.jpeg)

#### **Annihilation with atomic electrons**

annihilation process: 
$$e^+e_A^- \rightarrow A' - \sigma_{ann}(\sqrt{s}) = \frac{e^2\epsilon^2\alpha_D}{3} \frac{s+2m_\chi^2}{(s-m_{A'}^2)^2 + \Gamma_A^2/m_\chi^2}$$
  
 $\alpha_D = g_\chi^2/4\pi \qquad s = 2m_e E' + c_A^2/m_\chi^2$ 

![](_page_69_Figure_2.jpeg)

 $2m_e^2 = 2m_e E_{A'}$ 

![](_page_69_Figure_4.jpeg)

![](_page_69_Picture_5.jpeg)

## **Annihilation with atomic electrons (continued)**

$$N_{\rm ann} = \mathscr{L} \int_{E_{\rm min}}^{E_{\rm max}} dE \frac{d\sigma_B}{dE} \int_{0.95E}^{E+m_e} dE_{A'} n_e T_e$$

$$\frac{d\sigma_B}{dE}$$
 is the Bhabha xsec

 $n_{\rho}$  is the electron # density

 $T_{\rho}(E', E, L_T)$  is the  $e^+$  differential track length

[Tsai & Whitis 1966] [Bjorken et al, 1988]

 $\sigma_{e}(E' = E_{A'} - m_{e}, E, L_{T})\sigma_{ann}(E_{A'})$ 

![](_page_70_Figure_7.jpeg)

![](_page_70_Picture_9.jpeg)

#### **Bremsstrahlung with target nucleus**

#### dominated by on-shell A' production

$$N_{\rm bre} = \mathscr{L} \int_{E_{\rm min}}^{E_{\rm max}} dE \frac{d\sigma_B}{dE} \int_{0.95E}^{E-m_e} dE_{A'}$$

 $\frac{d\sigma_{\rm bre}}{dE_{A'}} = \operatorname{xsec} \text{ of on-shell produced } A'$ 

[Bjorken et al, 0906.0580] [Gninenko et al, 171205706] [Liu & Miller, 1705.01633]

 $A' n_N T_e(E', E, X_0) \frac{d\sigma_{\text{bre}}}{dE_{A'}}$ 

![](_page_71_Figure_7.jpeg)

![](_page_71_Picture_8.jpeg)

![](_page_71_Figure_9.jpeg)

![](_page_71_Figure_10.jpeg)
## Belle II sensitivity on invisible dark photon





We propose a new dark matter channel at colliders, where one SM particle interacts with the detector to produce DM particles

The main background at Belle II are due to photon and neutron events that escape the detection

invisible dark photon, surpassing both the mono-photon channel and NA64



We find that this new DM channel @ Belle II can probe new parameter space of



# backup slides



### **Track length**

For positrons with initial energy E to enter a target with thickness  $L_T$ , the differential track-length distribution as a function of the positron energy E' can be computed by [1, 2]

$$T_e(E', E, L_T) = X_0 \int_0^{L_T/X_0} I_e(E', E, t) dt,$$

where  $X_0$  is the radiation length of the target. Here  $I_e(E', E, t)$  is the energy distribution of E' at the depth  $tX_0$ , which can be computed iteratively such that  $I_e = \sum_i I_e^{(i)}$  where  $I_e^{(i)}$  denotes the *i*-th generation positrons [3]. We adopt the analytical model of Ref. [3] up to second-generation positrons, which are found to be in good agreement with simulations in Ref. [1]. The contributions from the first two generations are [3]

$$\begin{split} I_e^{(1)}(E',E,t) &= \frac{1}{E} \frac{(\ln(1/v))^{b_1 t - 1}}{\Gamma(b_1 t)}, \\ I_e^{(2)}(E',E,t) &= \frac{2}{E} \int_v^1 \frac{dx}{x^2} \frac{1}{b_2 + b_1 \ln(1 - x)} \left[ \frac{(1 - x)^{b_1 t} - (1 - v/x)^{b_1 t}}{b_1 \ln\left[(x - x^2)/(x - v)\right]} + \frac{e^{-b_2 t} - (1 - v/x)^{b_1 t}}{b_2 + b_1 \ln(1 - v/x)} \right], \end{split}$$

where  $b_1 = 4/3$ ,  $b_2 = 7/9$ , v = E'/E.

[1] 1802.03794 [2] 1807.05884 [3] Tsai & Whitis 1966













### xsec of on-shell dark photon

where  $n_N$  is the number density of I (or Cs). Here  $d\sigma_{\rm bre}/dE_{A'}$  is the differential cross section of the on-shell produced A' [71–73],

$$\frac{d\sigma_{\rm bre}}{dE_{A'}} = (\phi_I + \phi_{\rm Cs}) \frac{4\alpha^3 \epsilon^2}{E'} \frac{x(1 - x + x^2/3)}{m_{A'}^2(1 - x) + m_e^2 x^2}, \quad (13)$$

where  $x \equiv E_{A'}/E'$ , and  $\phi_N$  denotes the effective flux of photons from nucleus N [71]:

$$\phi_N = \int_{t_{\min}}^{t_{\max}} dt \, \frac{t - t_{\min}}{t^2} \left[ \frac{Za^2 t}{(1 + a^2 t)(1 + t/d)} \right]^2, \quad (14)$$

with  $t_{\min} = (m_{A'}^2/2E')^2$ ,  $t_{\max} = m_{A'}^2 + m_e^2$ ,  $a = 111m_e^{-1}Z^{-1/3}$ , and  $d = 0.164A^{-2/3}$  GeV<sup>2</sup>. We use Z = 53(55) and A = 127(133) for I (Cs). Here we only consider the dominant elastic form factor.

[71] Bjorken et al, 0906.0580 [72] Gninenko et al, 171205706 [73] Liu & Miller, 1705.01633

